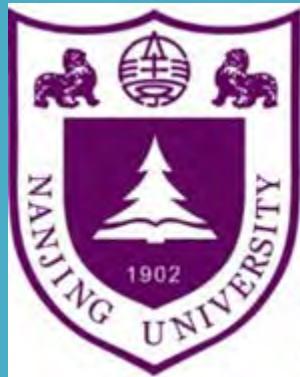


# Summary of the **Solar/Astron Session**

**P. F. Chen**



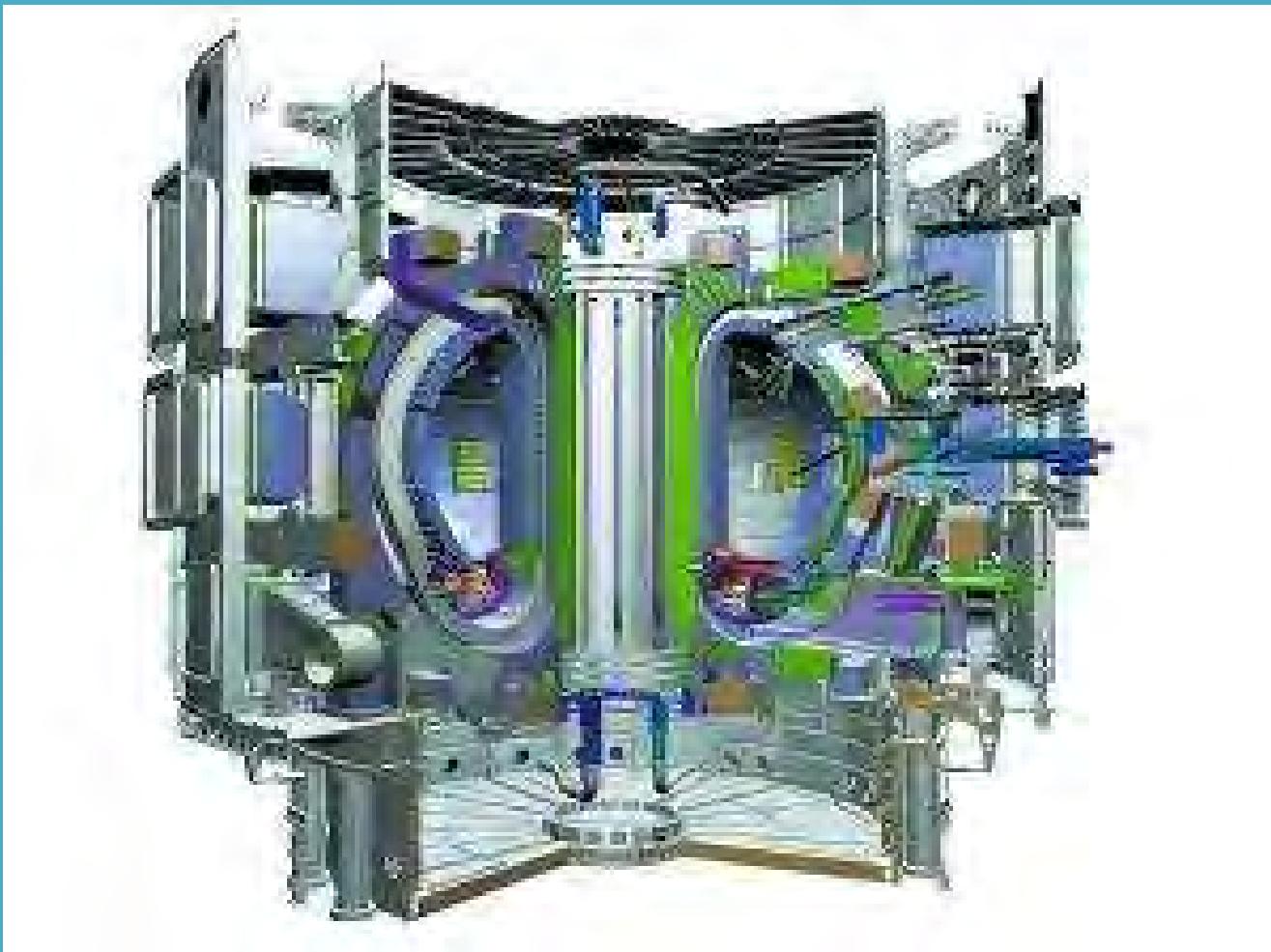
**[chenpf@nju.edu.cn](mailto:chenpf@nju.edu.cn)**

**NANJING UNIVERSITY**



## Charm of *Chengdu*

Once being here,  
you would rather **live** here,  
not to **leave** here.



Solar/Astron guys love instabilities!

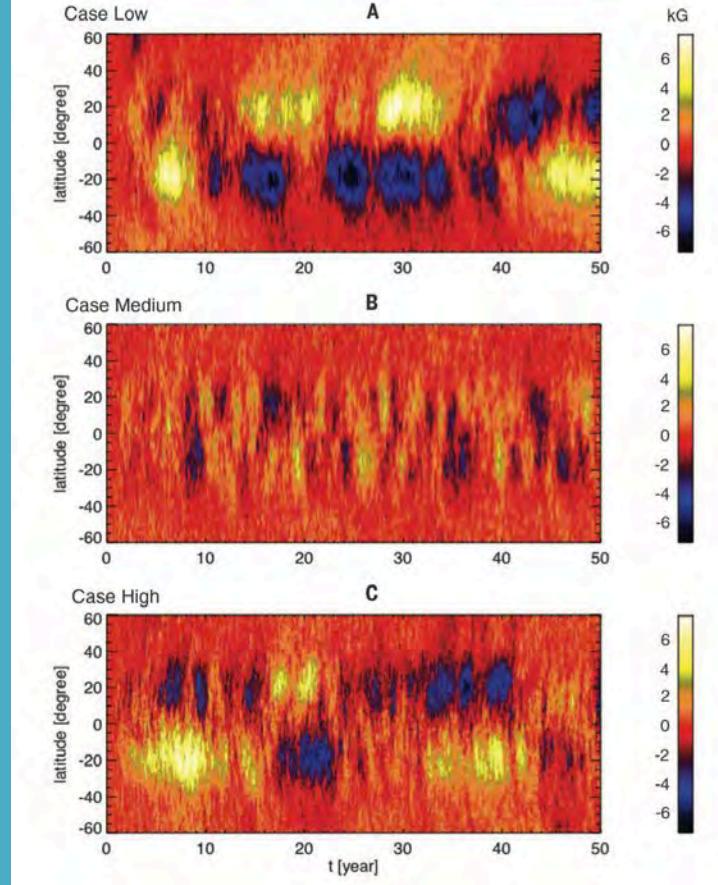
# Dynamo, Helicity & Filaments



M. Zhang

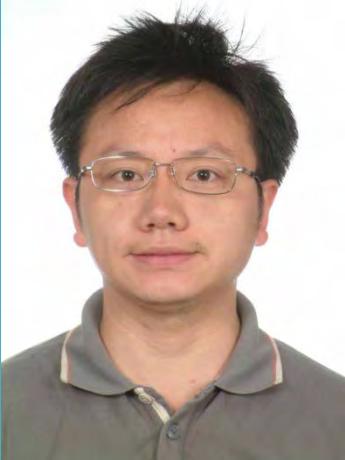


J. Z. Zhu



Hotta +  
(2016)

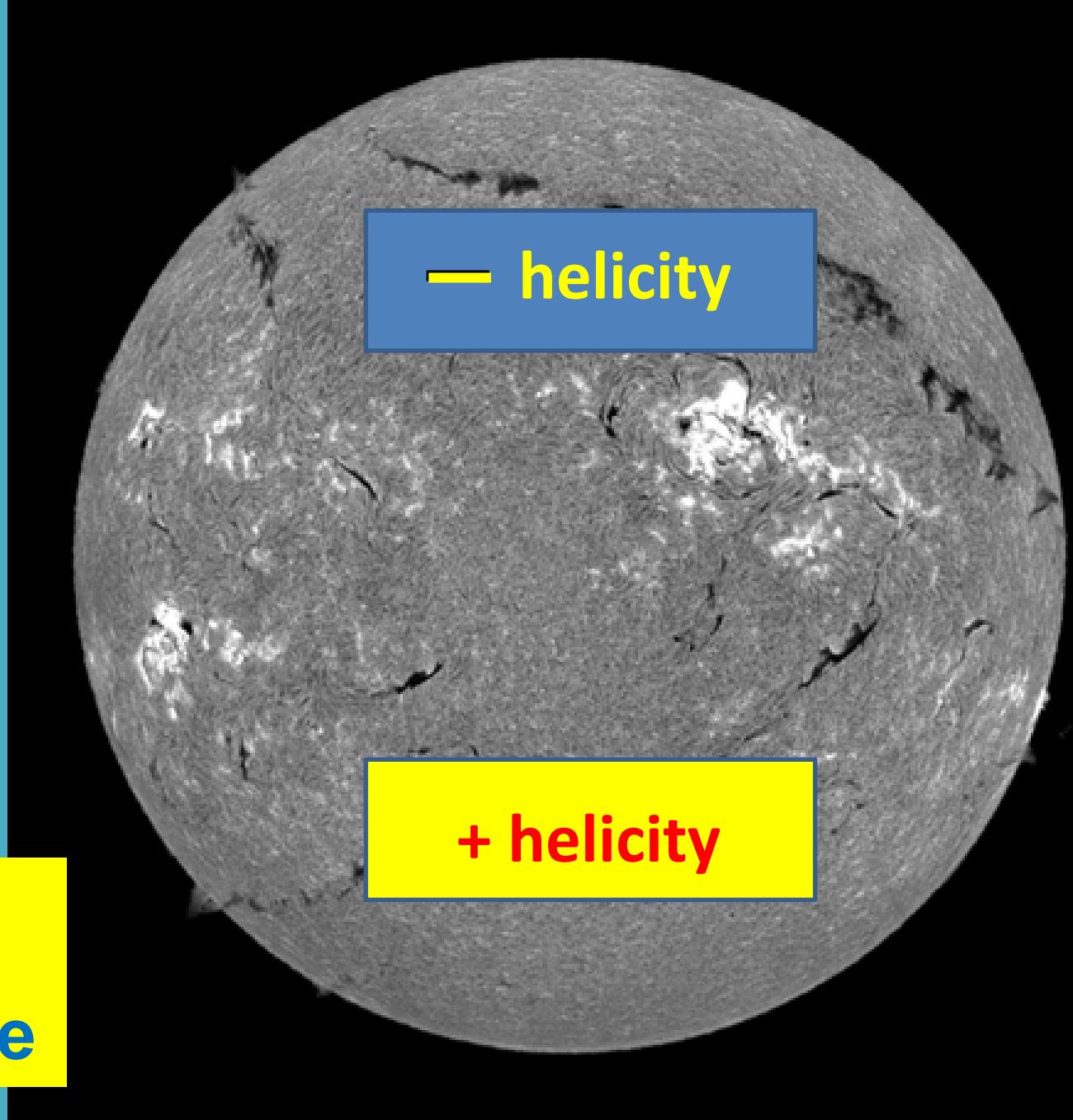
$$\alpha = \frac{\tau_c}{3} \left( -\overline{\boldsymbol{\omega} \cdot \boldsymbol{u}} + \frac{\overline{\boldsymbol{j} \cdot \boldsymbol{b}}}{\overline{\rho}} \right)$$

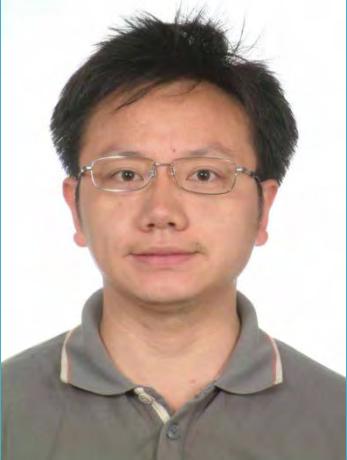


P. F. Chen

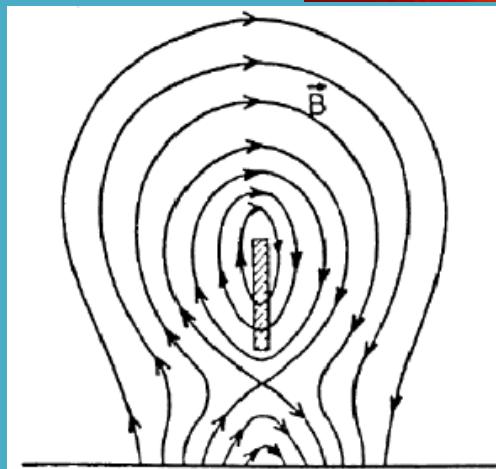
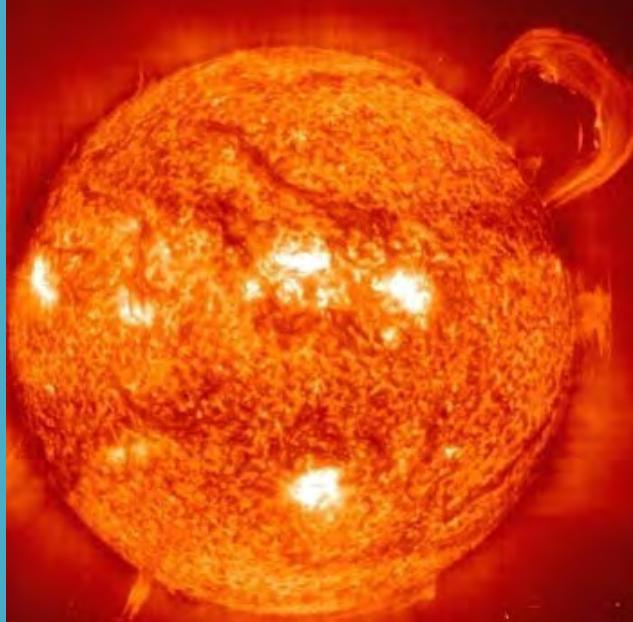
$$H_c = \mathbf{J} \bullet \mathbf{B}$$

**91.6%**  
**follow the rule**



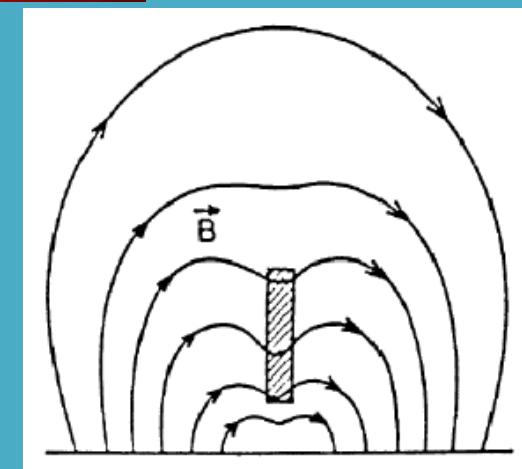


P. F. Chen



Flux rope

89%



Sheared arcade

11%



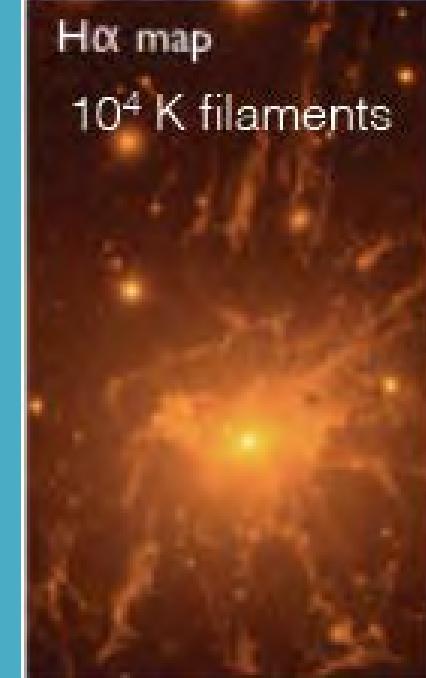
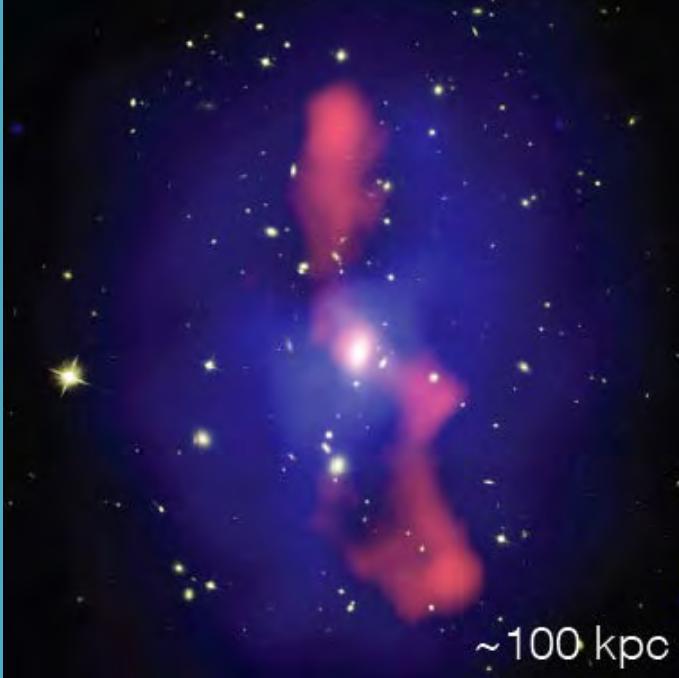
A. Hillier



Magnetic Rayleigh-Taylor instability



P. Sharmar



Mass supply cycle:

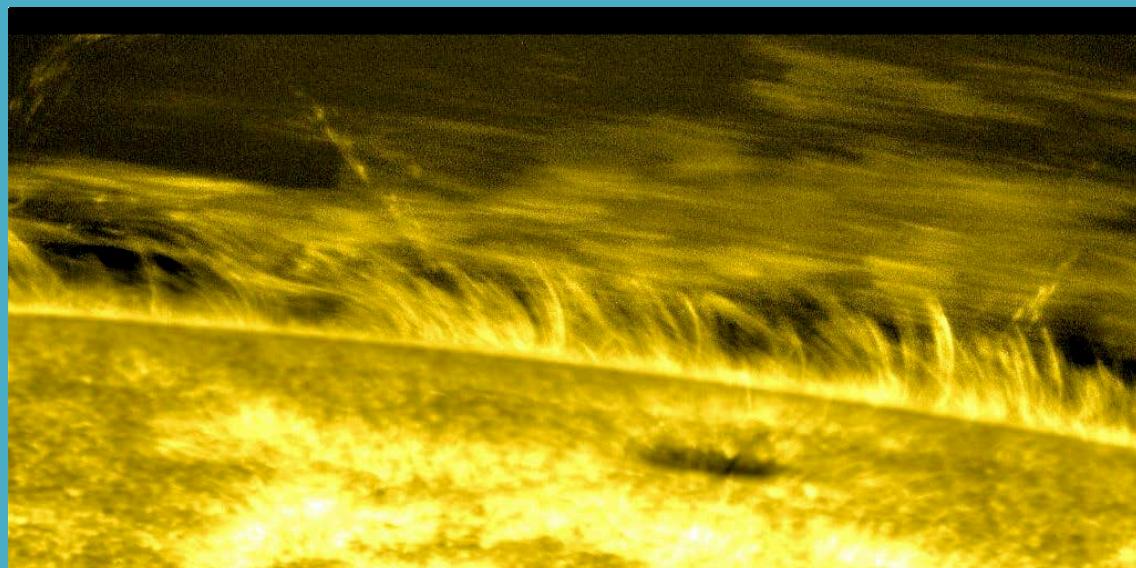
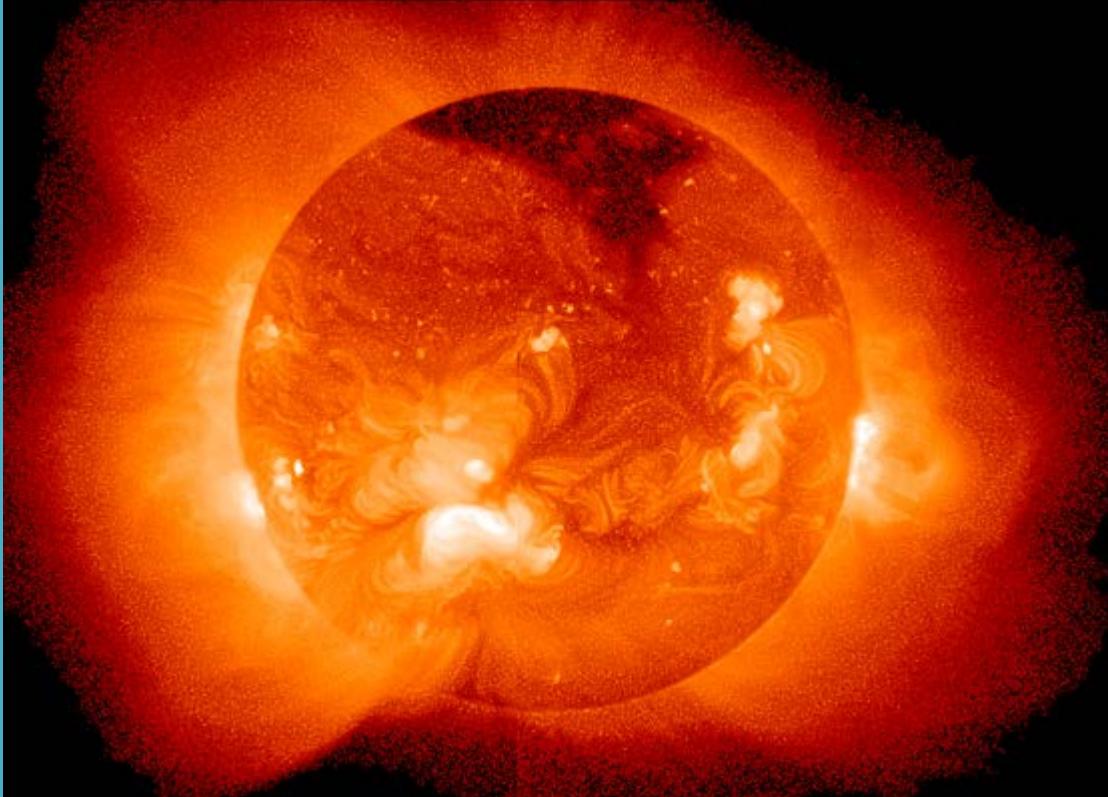
Black hole  $\rightarrow$  Corona  $\rightarrow$  Cool filaments



# Coronal Heating



J. Okamoto



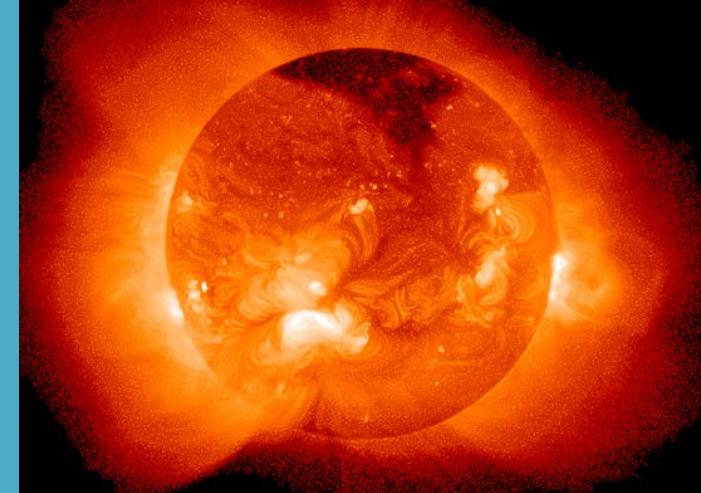
Alfven waves



Ling Chen

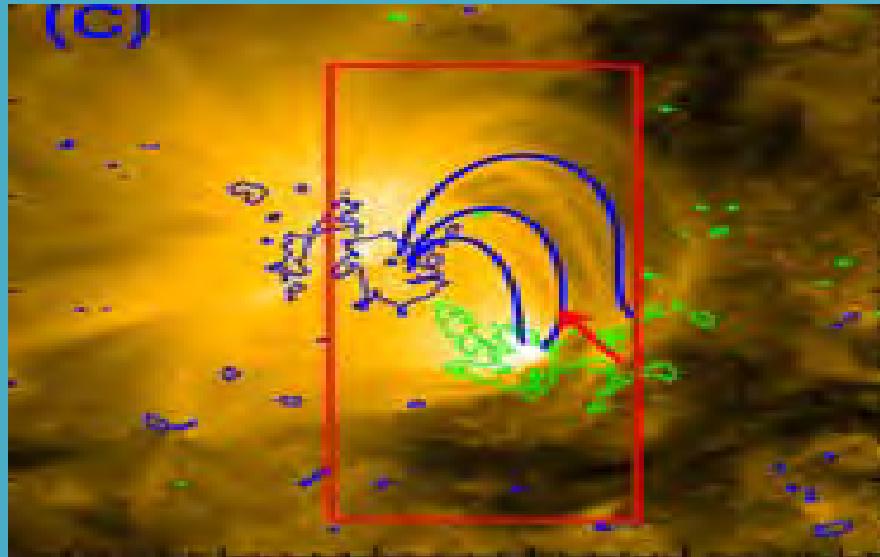


L. Xiang



## Kinetic Alfvén waves via Landau damping

Ion beams are more effective than electron beams in exciting Alfvén cyclotron waves.



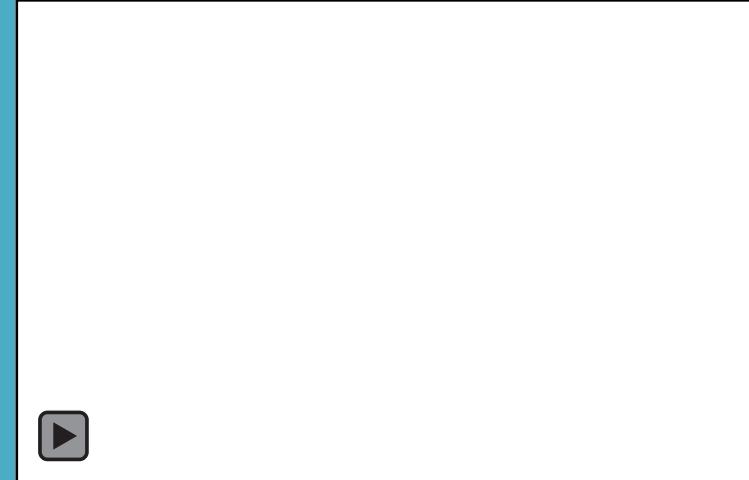
H. X. Xie

Most loops are not in a hydrostatic state

T is nearly uniform along the loops.

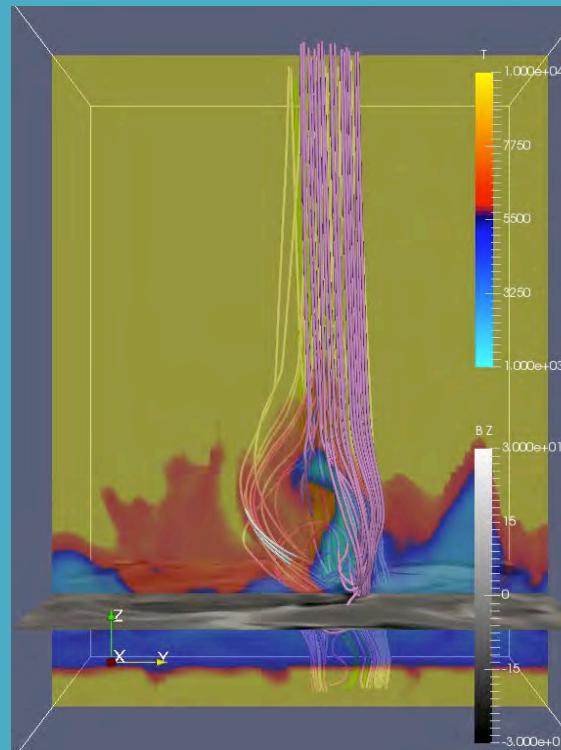


# Solar chromospheric jets



T. Yokoyama

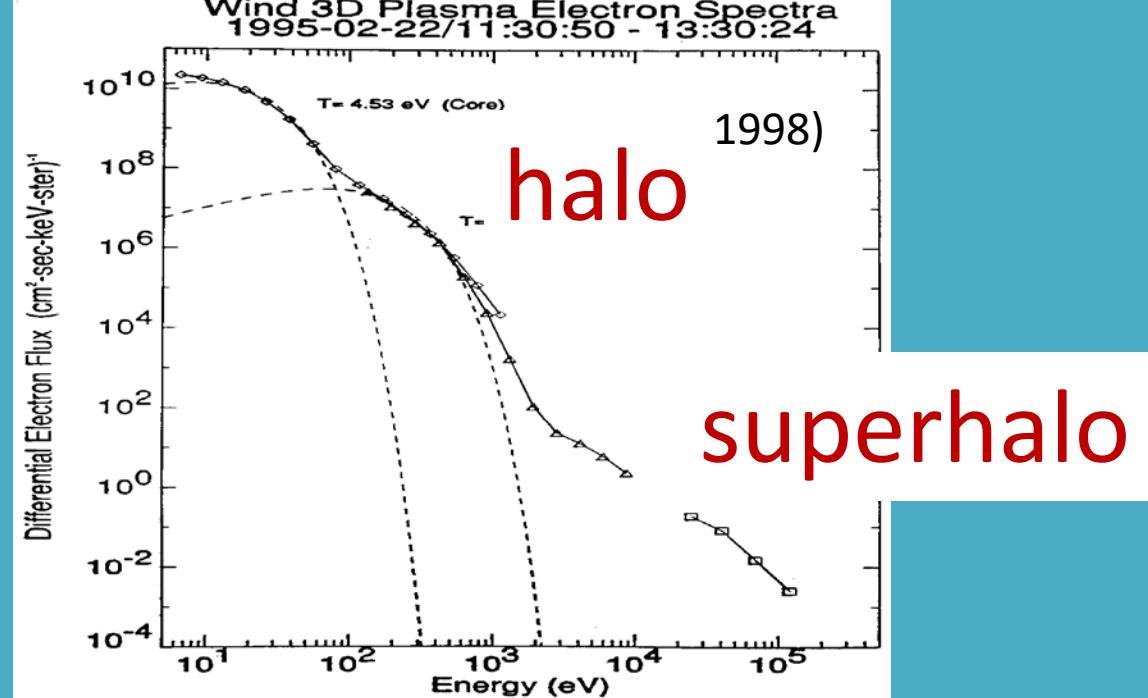
3D MHD



# Solar Wind



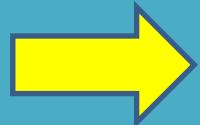
L. H. Wang



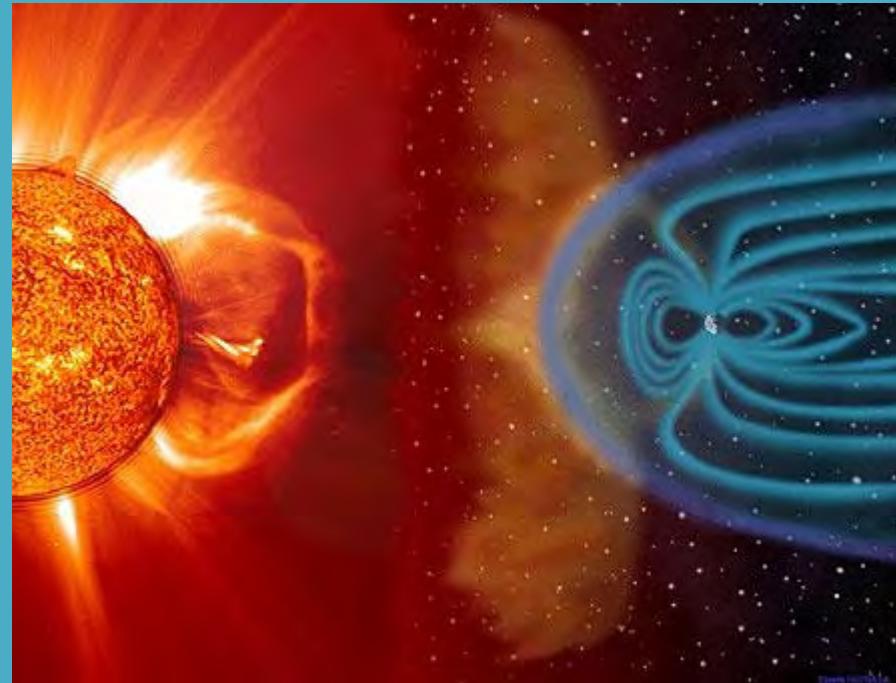
A power-law spectrum,  $J \sim E^{-\beta}$ , with  $\beta \sim 2.4$ .

Nearly isotropic angular distributions.

No clear correlation with solar flares or CMEs.



Different origins



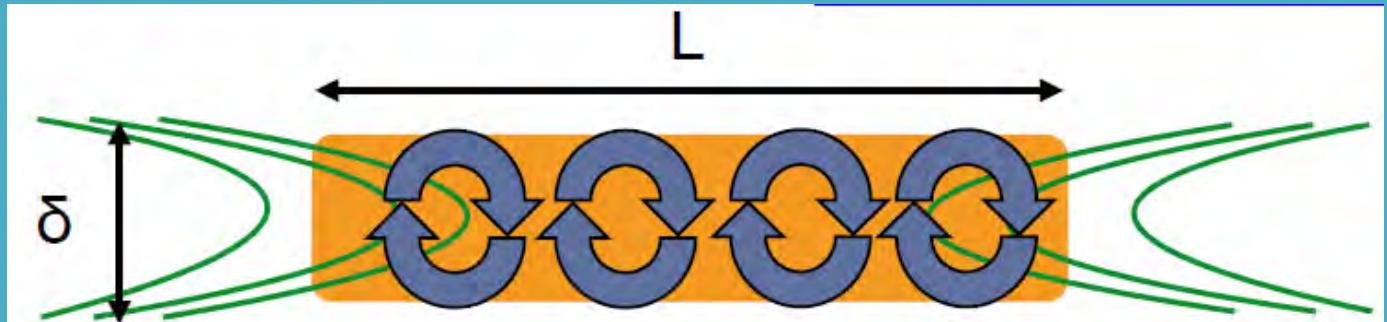
J. Seough

Whistler waves can be generated by electrons.

# Relativistic MHD

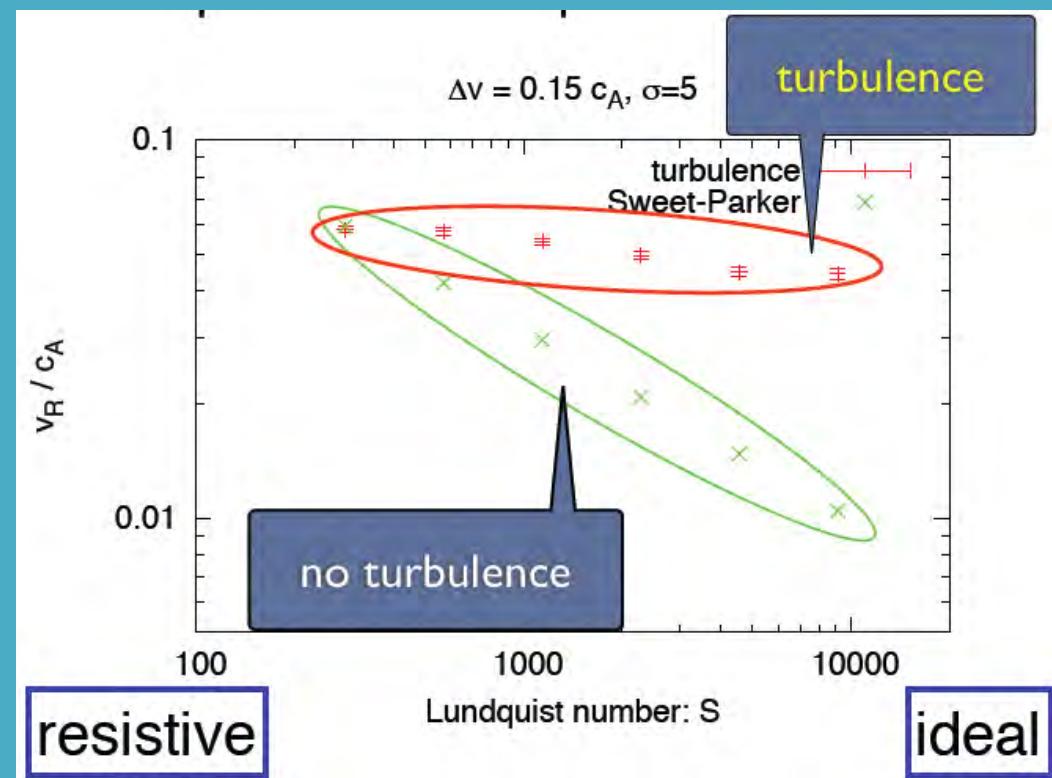


# Relativistic reconnection in Poynting-dominated plasma



M. Takamoto

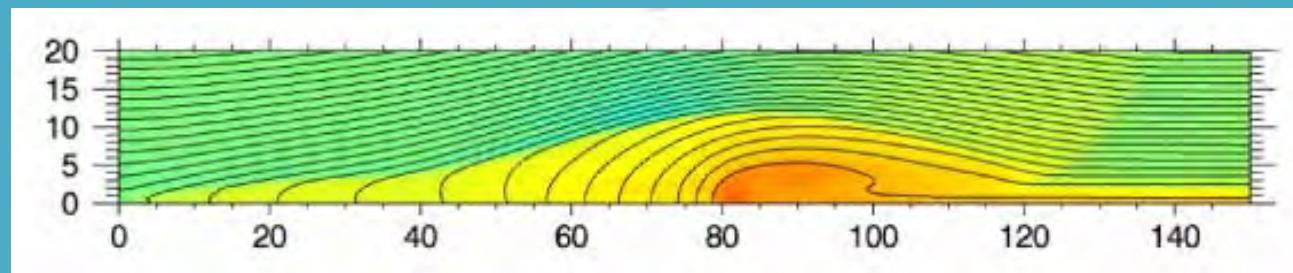
Only 1% of B energy  
is sufficient to drive  
turbulent reconnection





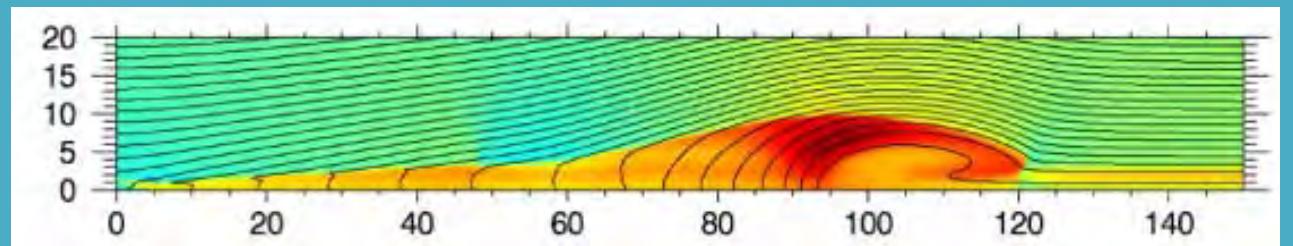
# Radiative relativistic reconnection

without radiation



T. Takeshige

With radiation

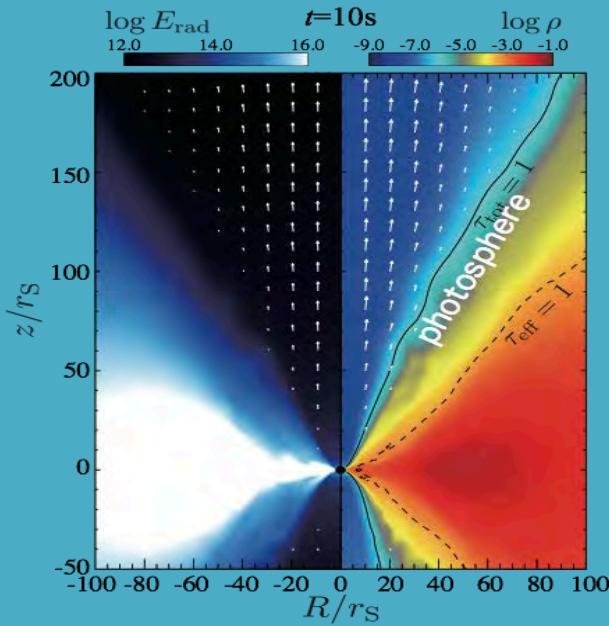


Radiation can increase reconnection rate.

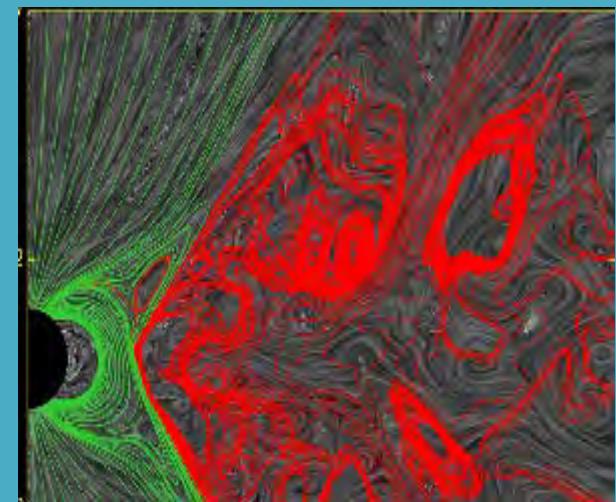


# Radiative MHD with GR

Black hole



Neutron star

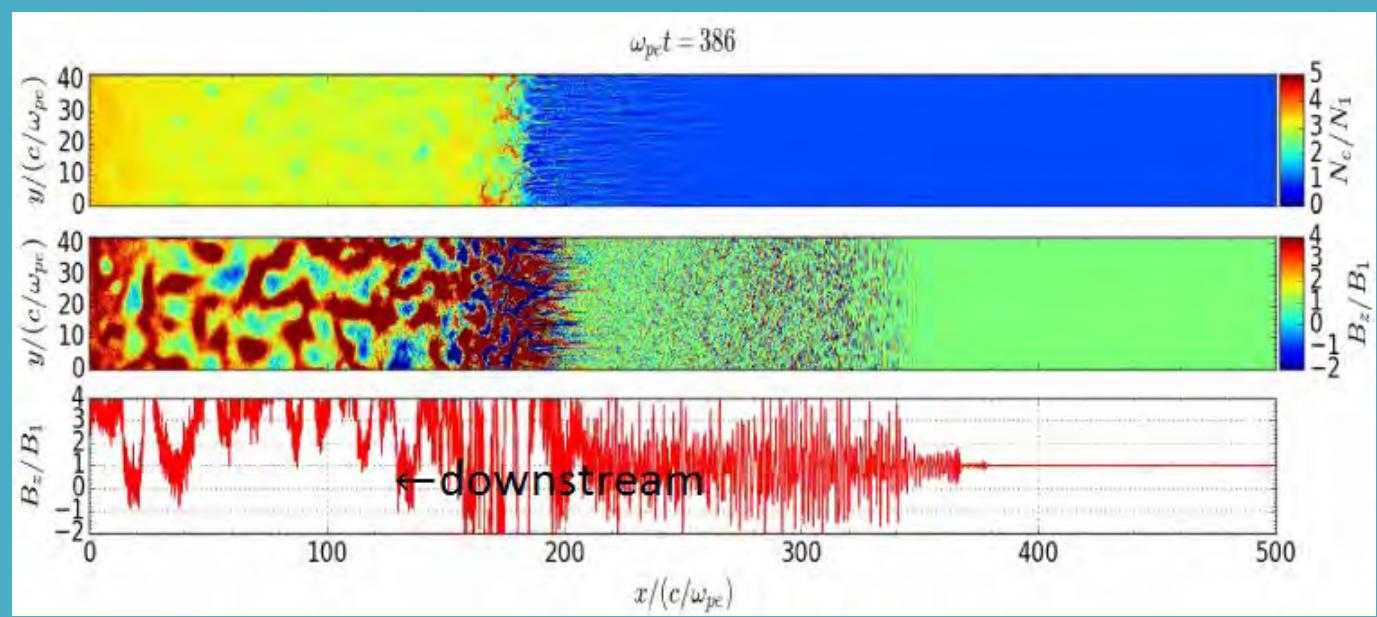


H. Takahashi

Radiation can accelerate high-speed jets



M. Iwamoto  
PIC  
simulations



Strong electromagnetic  
wave emissions are  
generated at the upstream  
of relativistic shocks

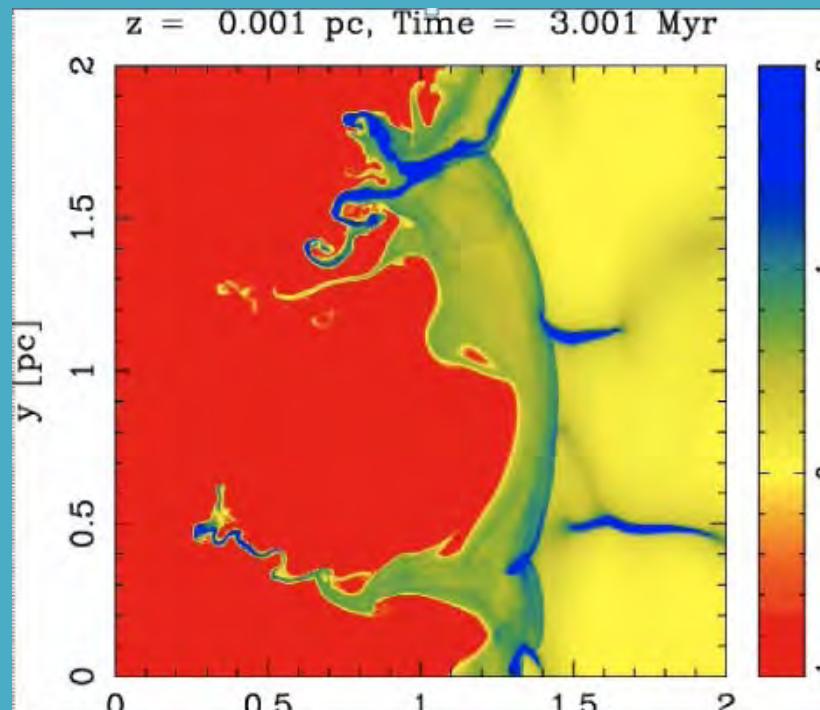
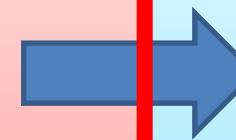


T. Inoue

Uniformity is boring.

Downstream

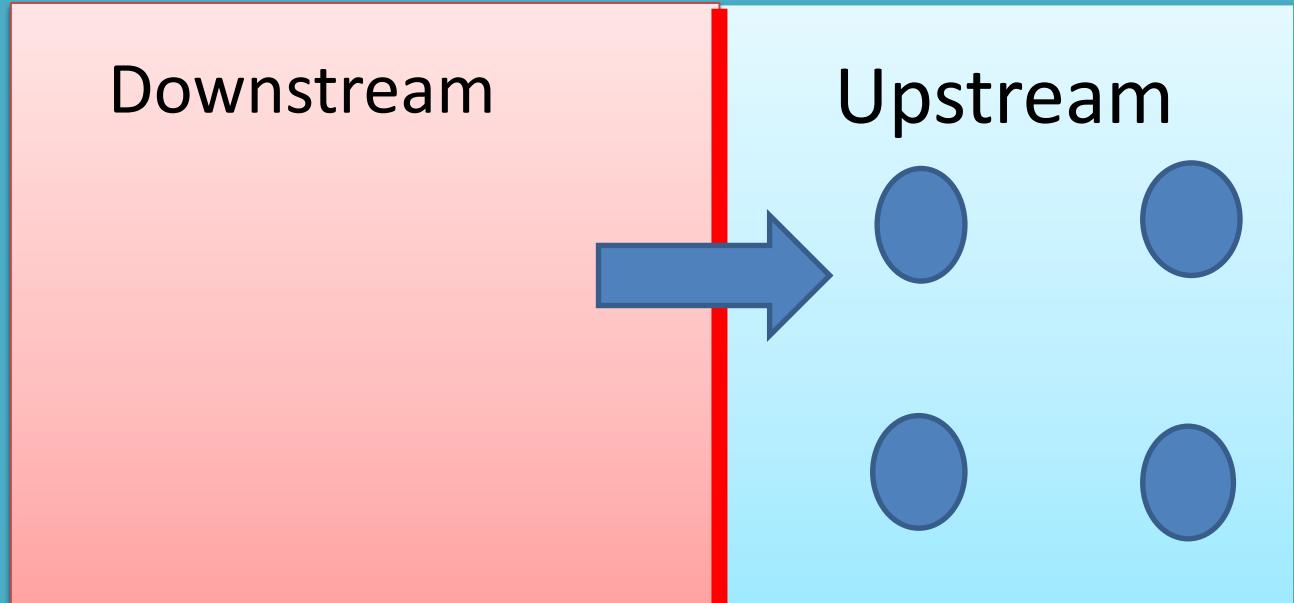
Upstream



Clumps help accelerate particles



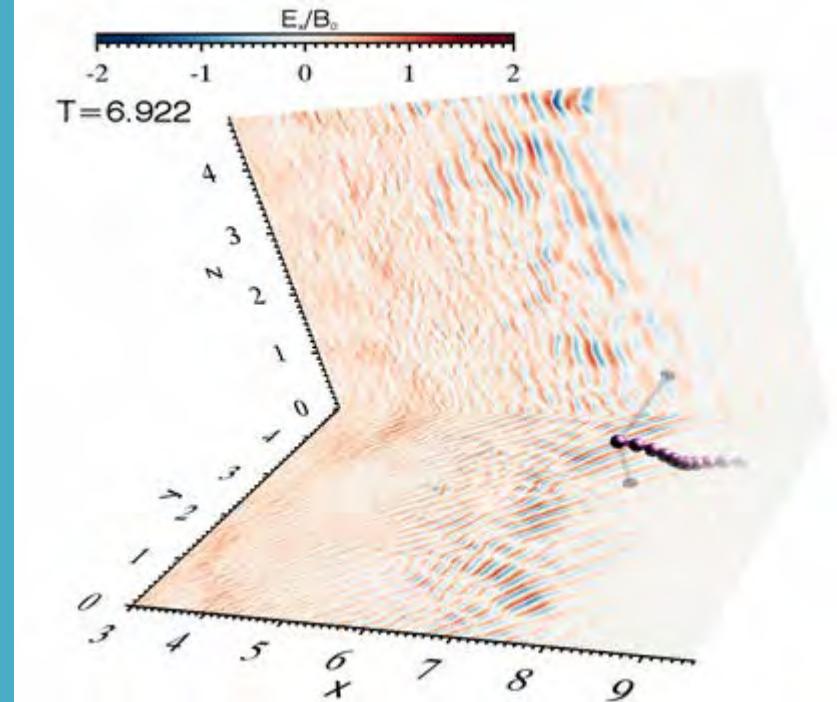
S. Tomita



Non-uniform upstream can extend the lifetime of the enhanced B in the downstream.

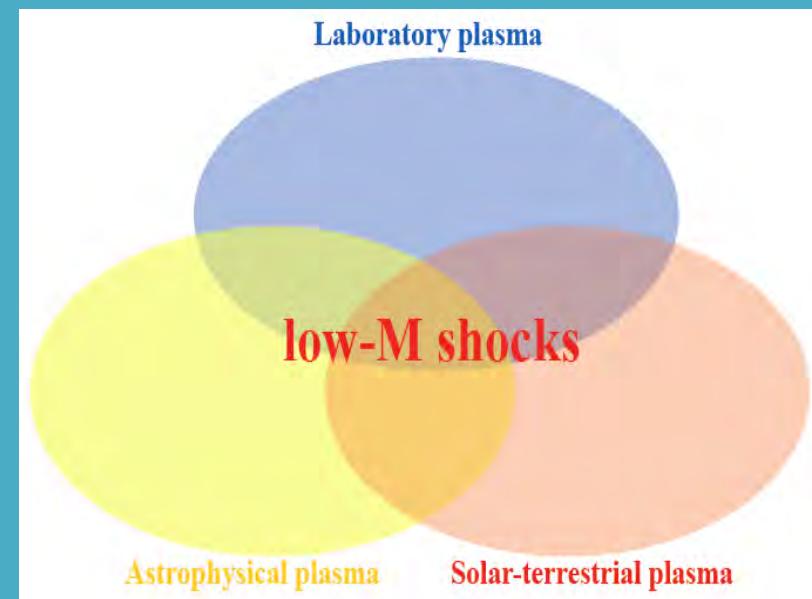
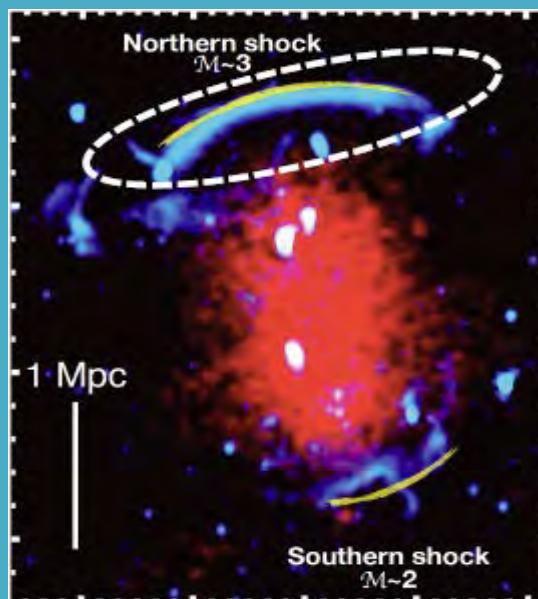


Y. Matsumoto



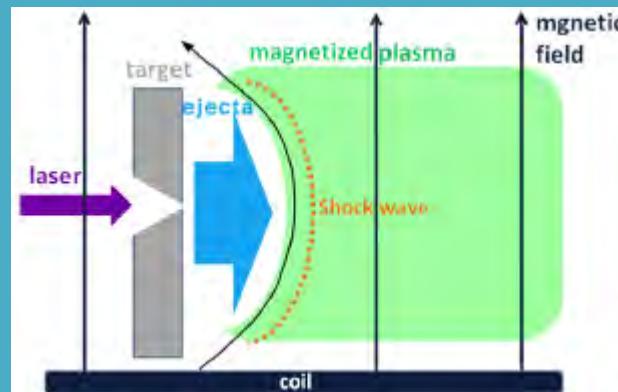
The 1<sup>st</sup> 3D PIC simulations of high- $M_A$  shock

Electron shock surfing & drift accelerations  
under turbulence generate relativistic particles



R. Yamazaki

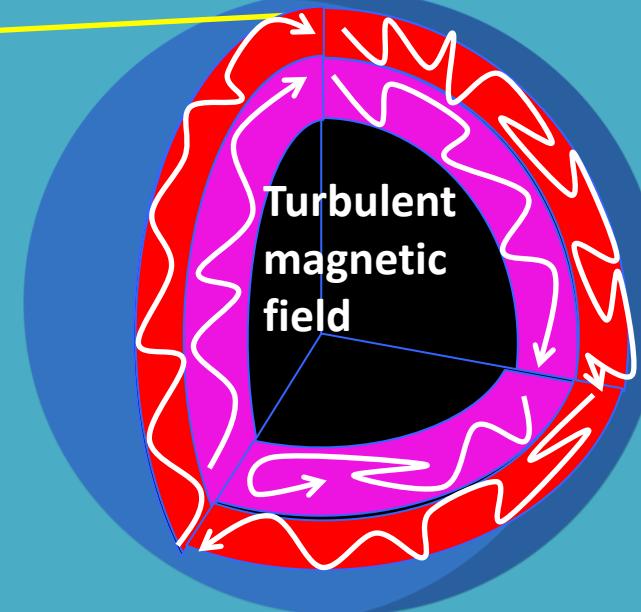
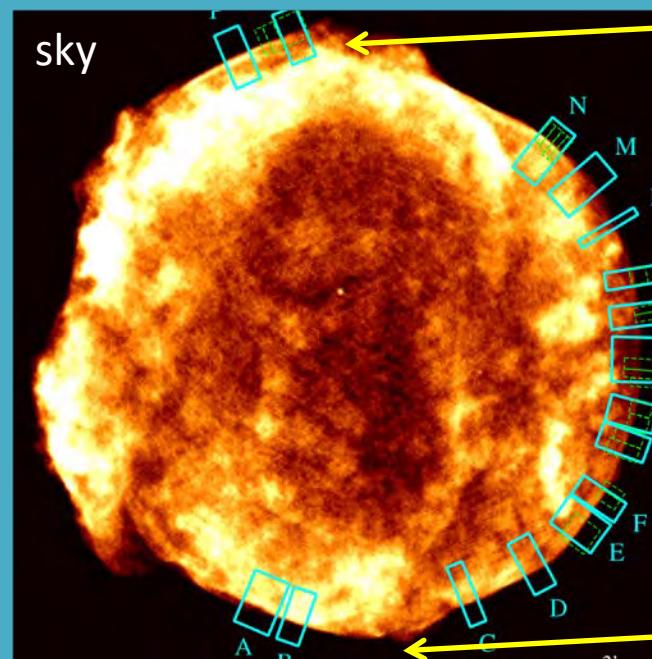
PIC simulations +



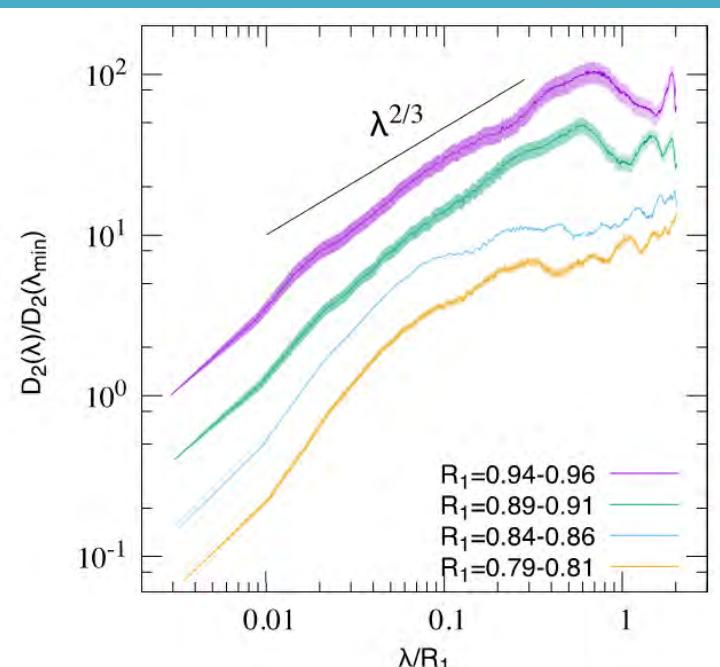
**Advantage:** simultaneous measurements of density and T across the shock,  $\text{Ma}>3$



J. Shimoda



B → shock → particle → radio emission

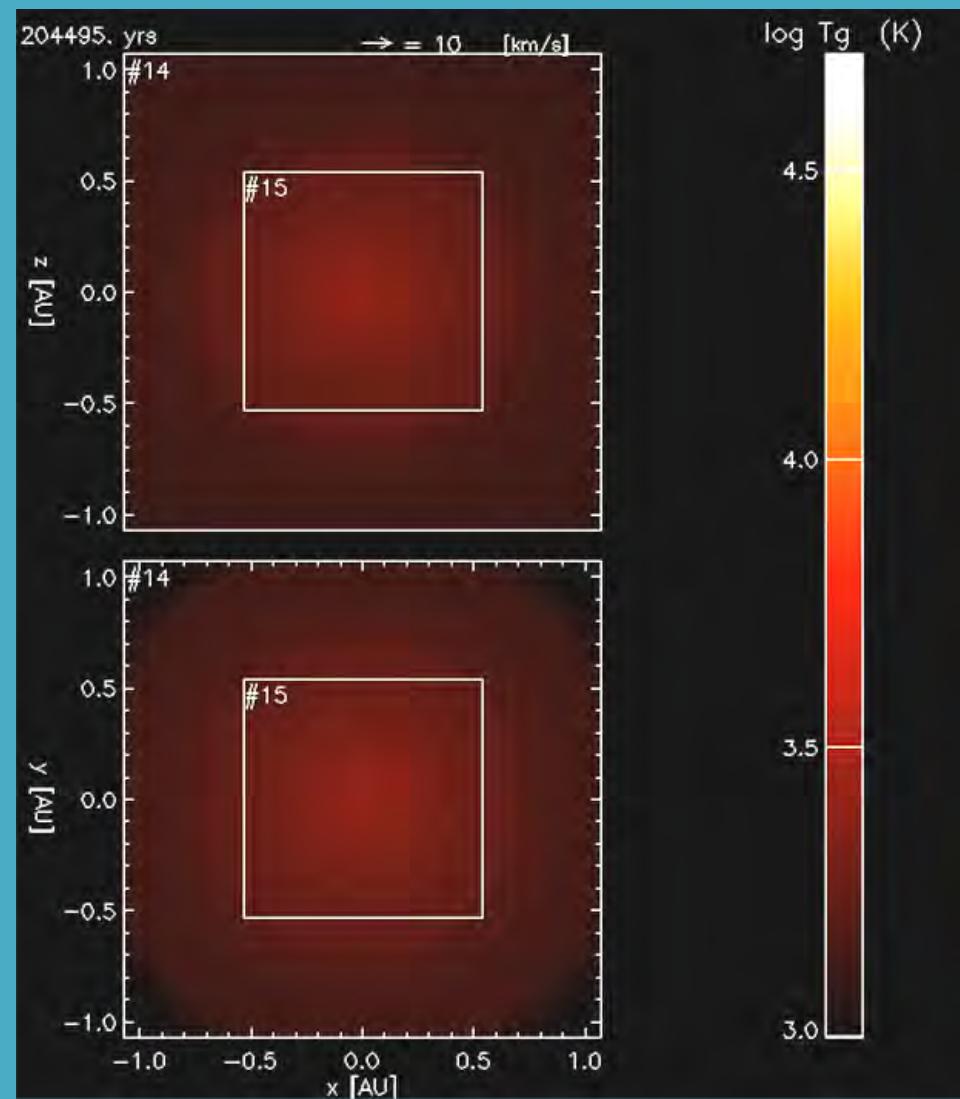


Komolgorov-law only  
at the outer shell

# Star Formation



K. Tomida



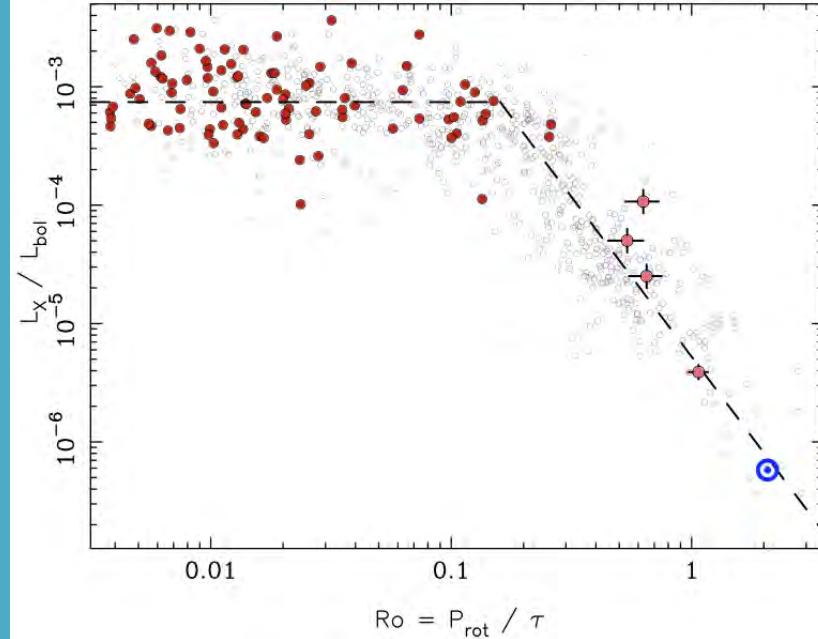
Sun and planets are formed together



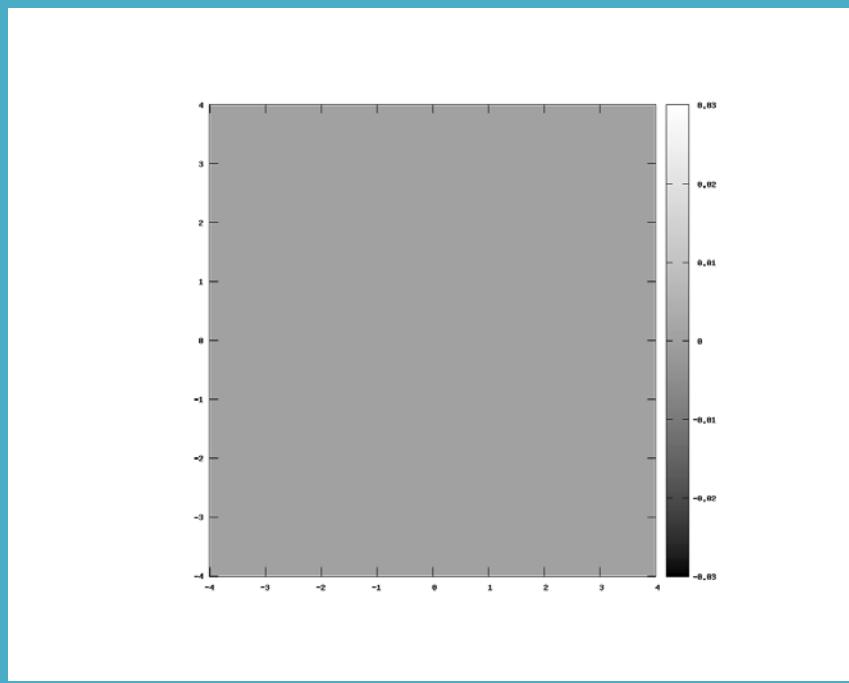
Y. Masada

A simple MHD  
model to explain it

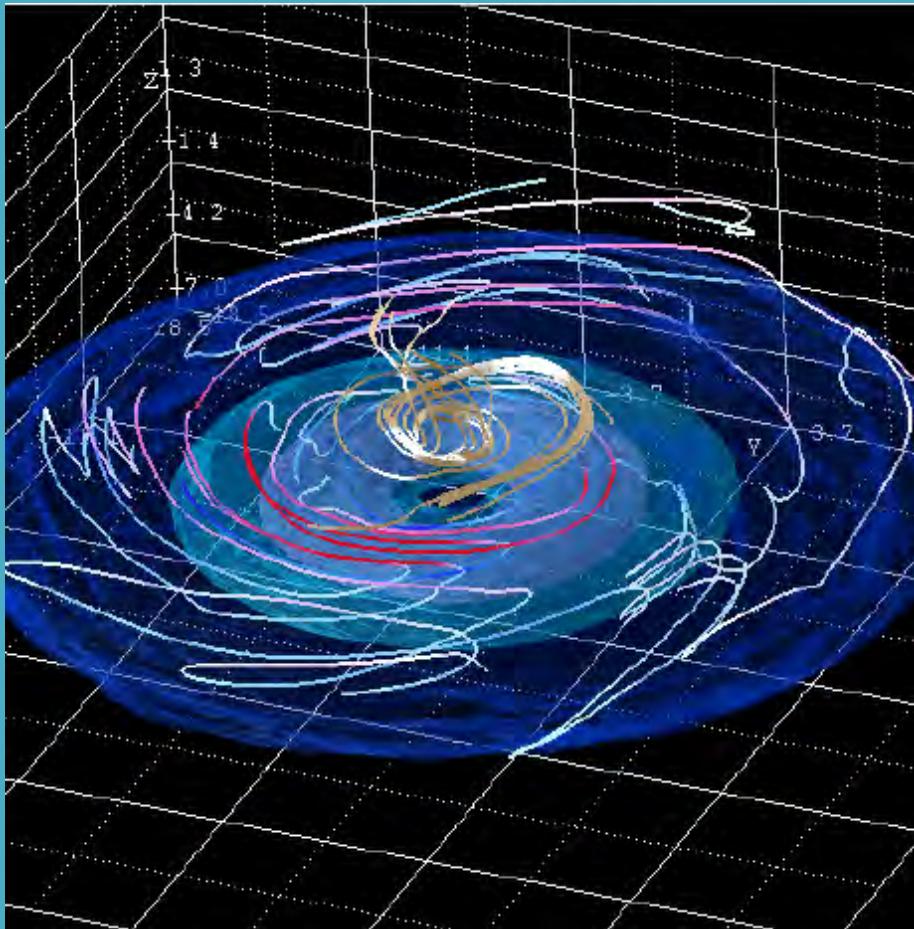
11-year solar cycle

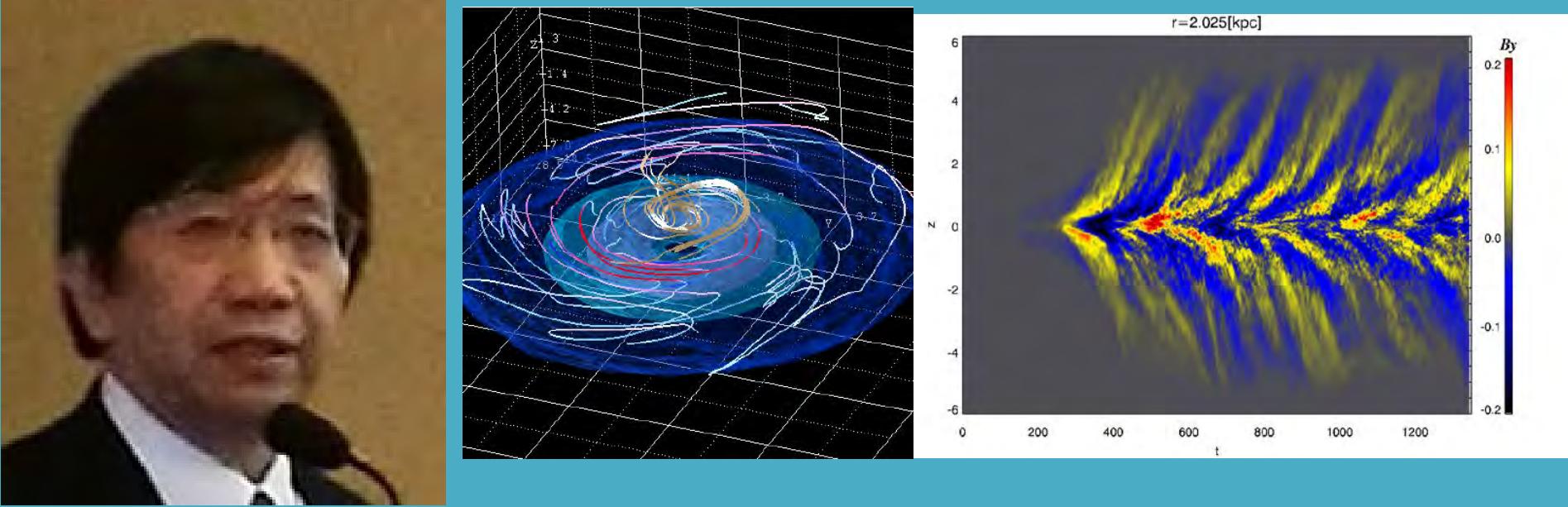


A Key Parameter : Rossby number



# Accretion Disk

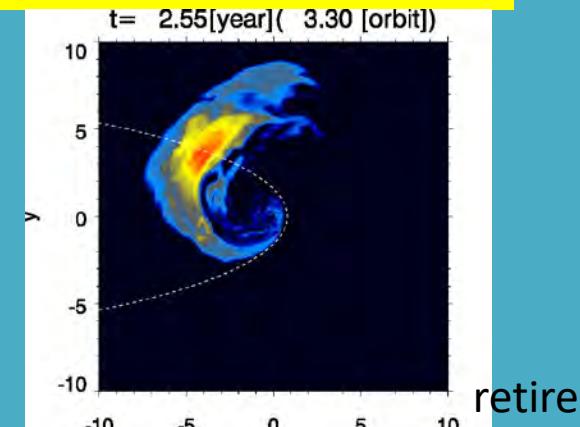
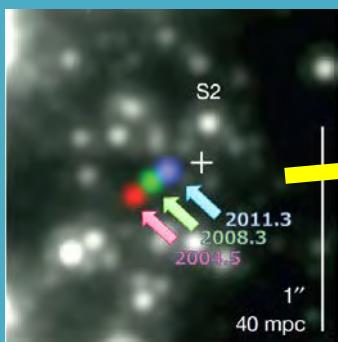




R. Matsumoto

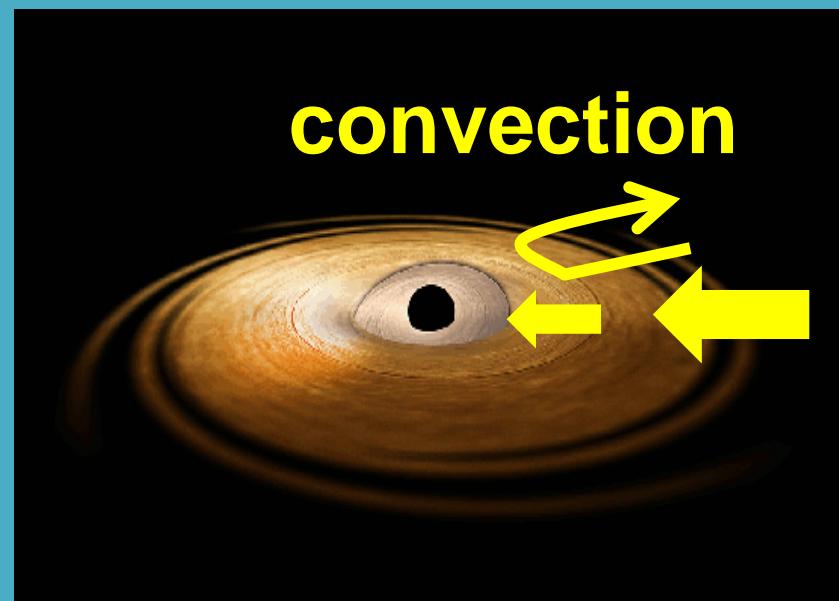
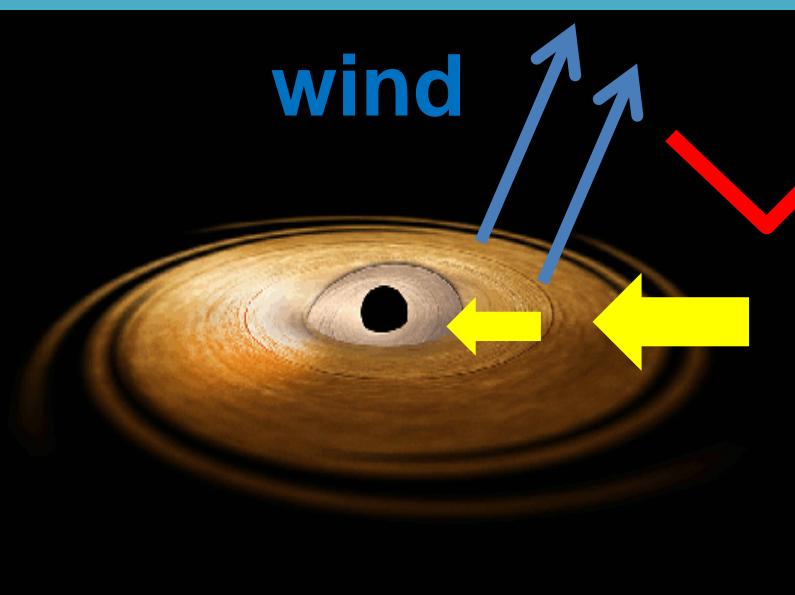
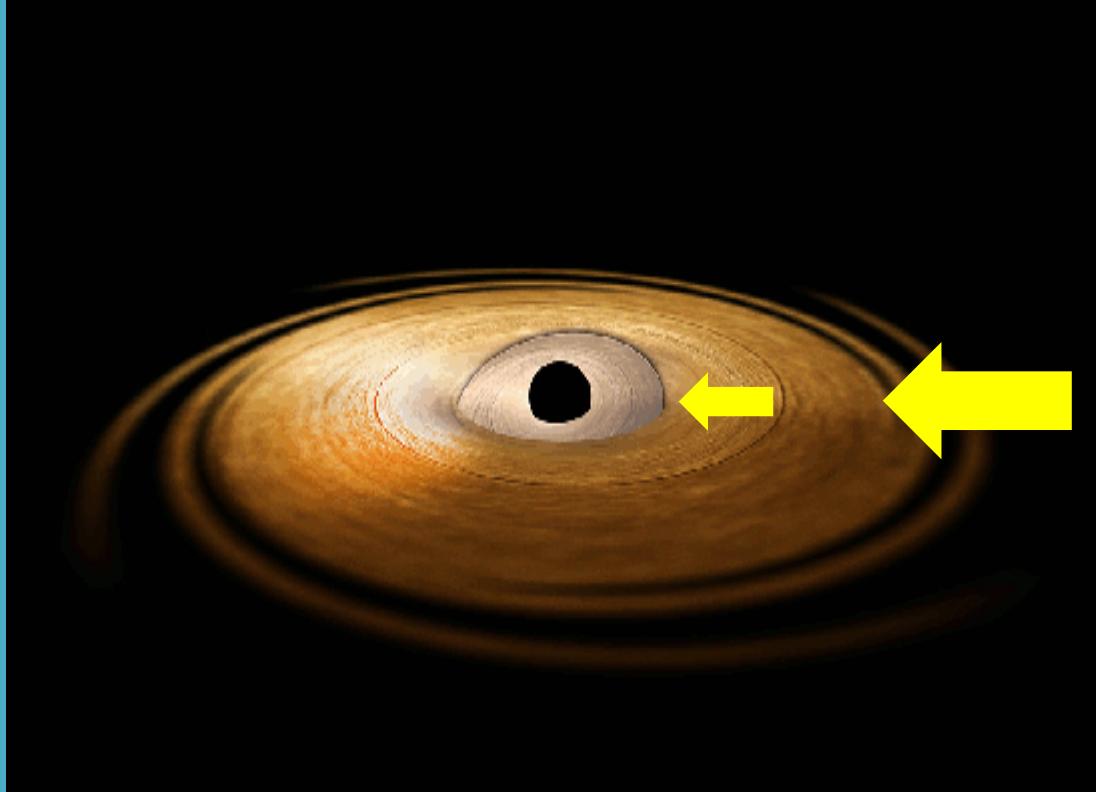
A higher order MHD code CANS+

MRI + buoyant escape → cycle





D. F. Bu



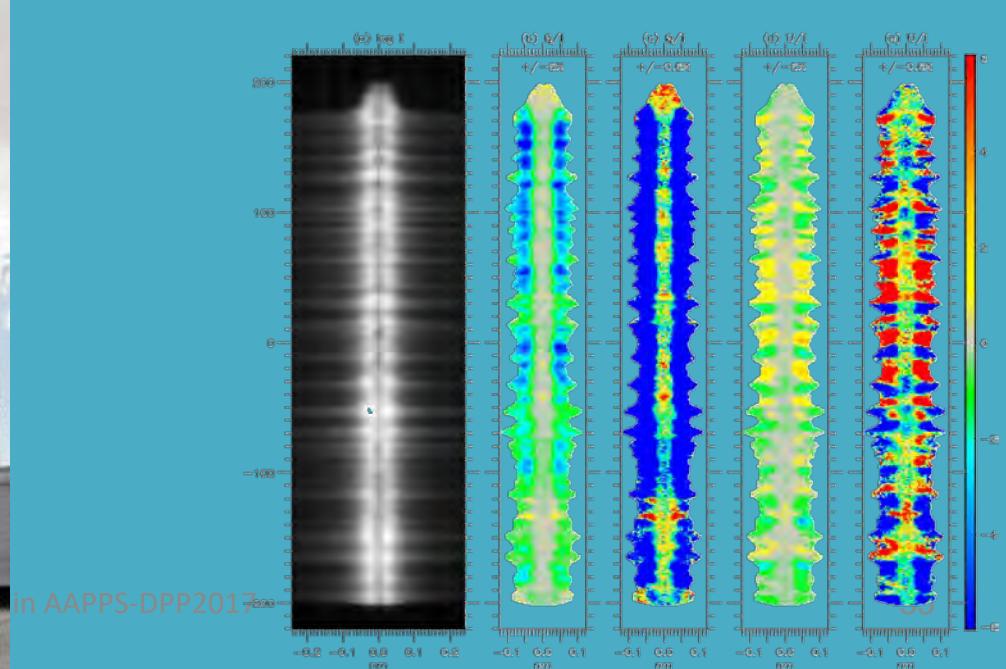
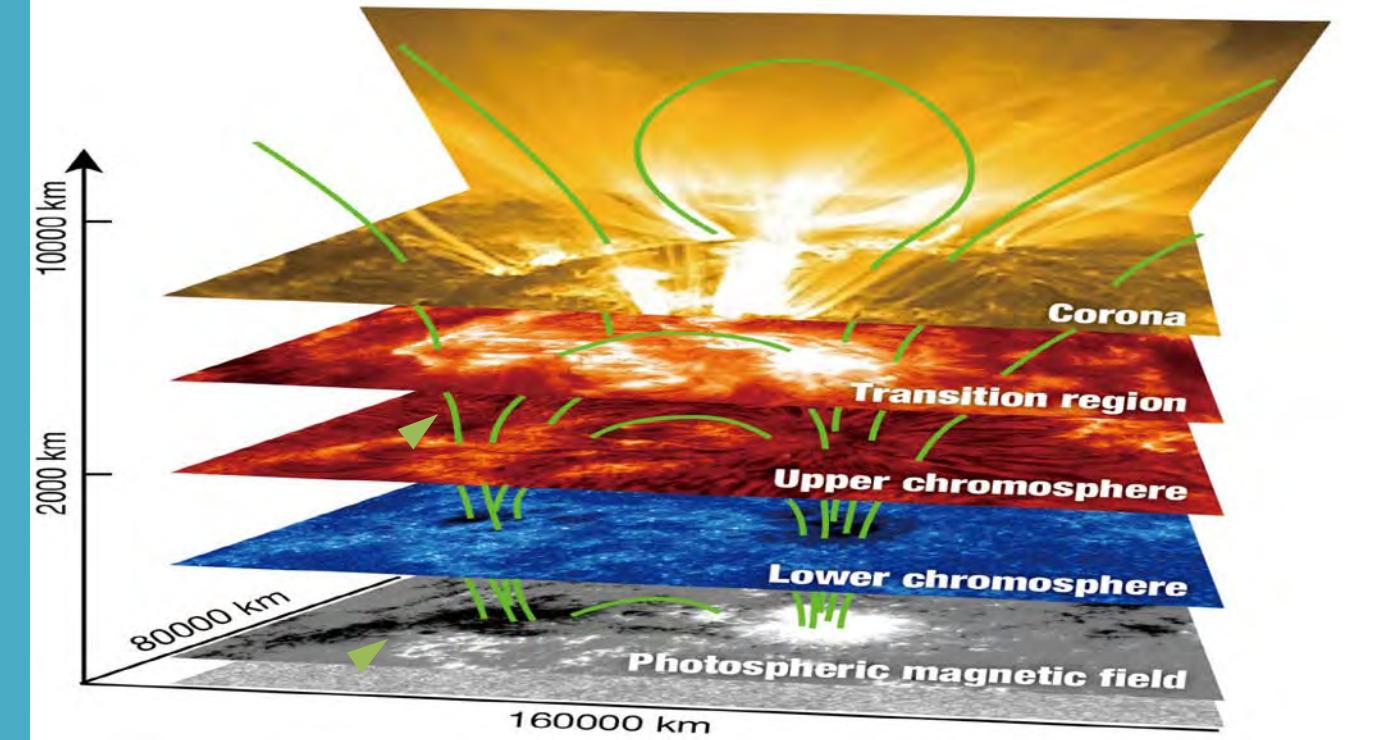
# Instruments



# R. Kano

## CLAPS

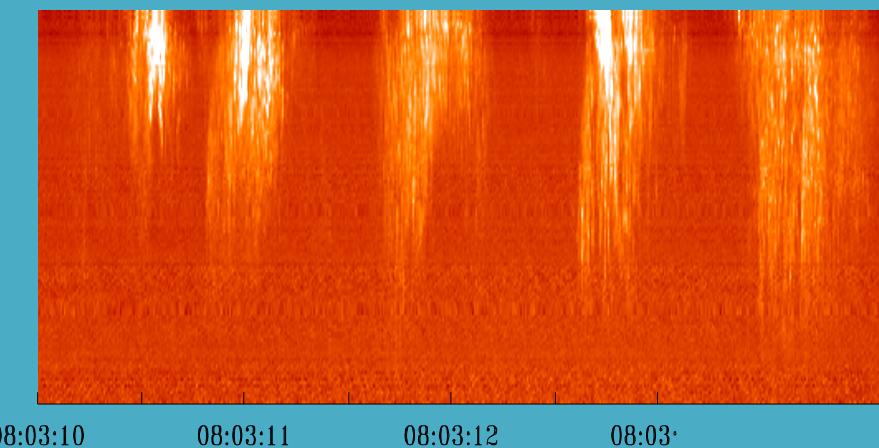
2017/09/18



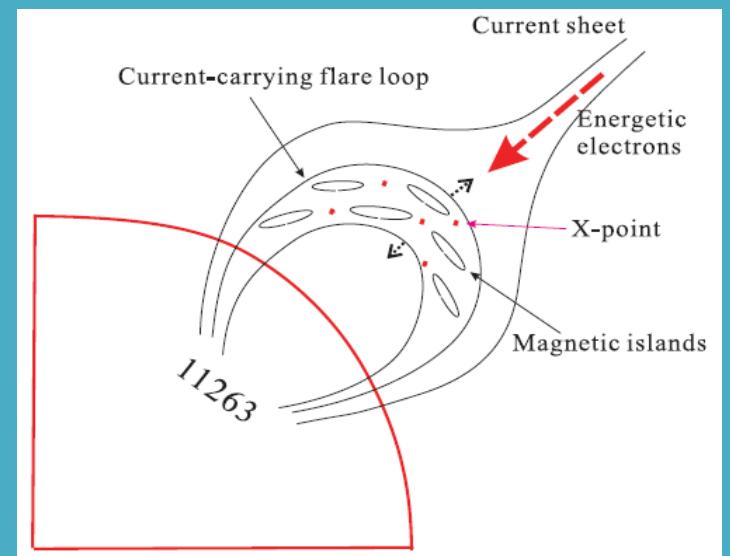
in AAPPS-DPP2017

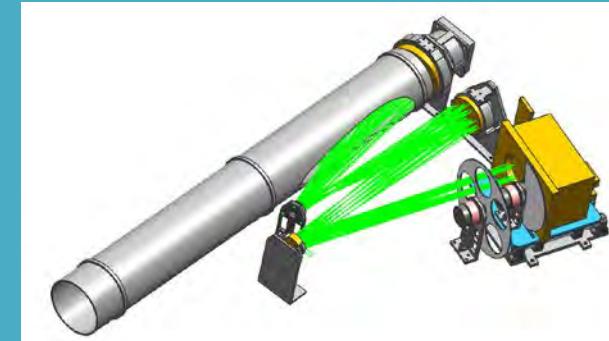


B. L. Tan



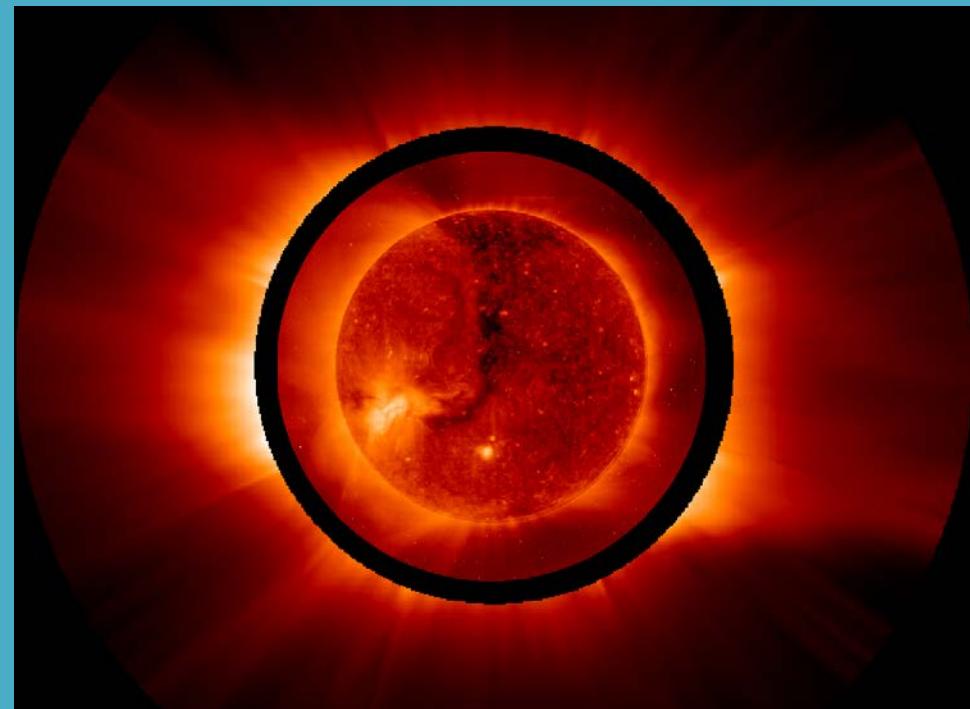
Mingantu Spectral Radioheliograph—MUSER  
100 antenna, max. baseline: 3 km, array with 3 spiral arms





H. Li

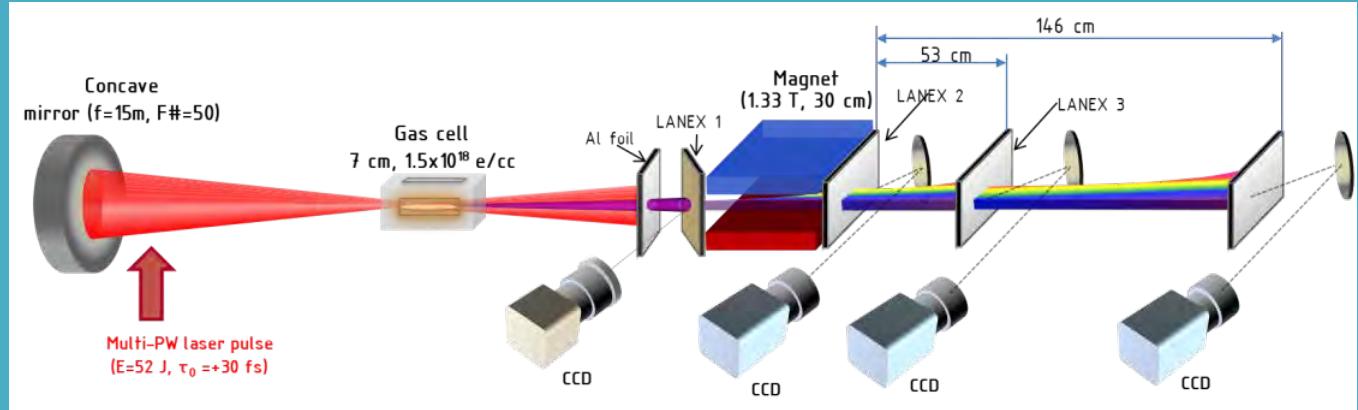
Ly-alpha  
Solar Telescope  
(LST)



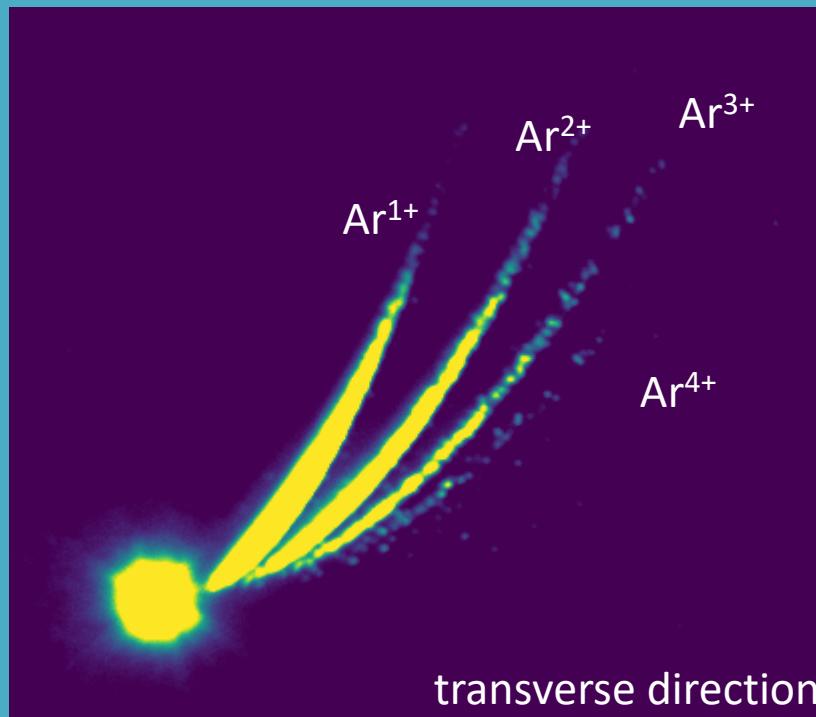
# Lab Experiments

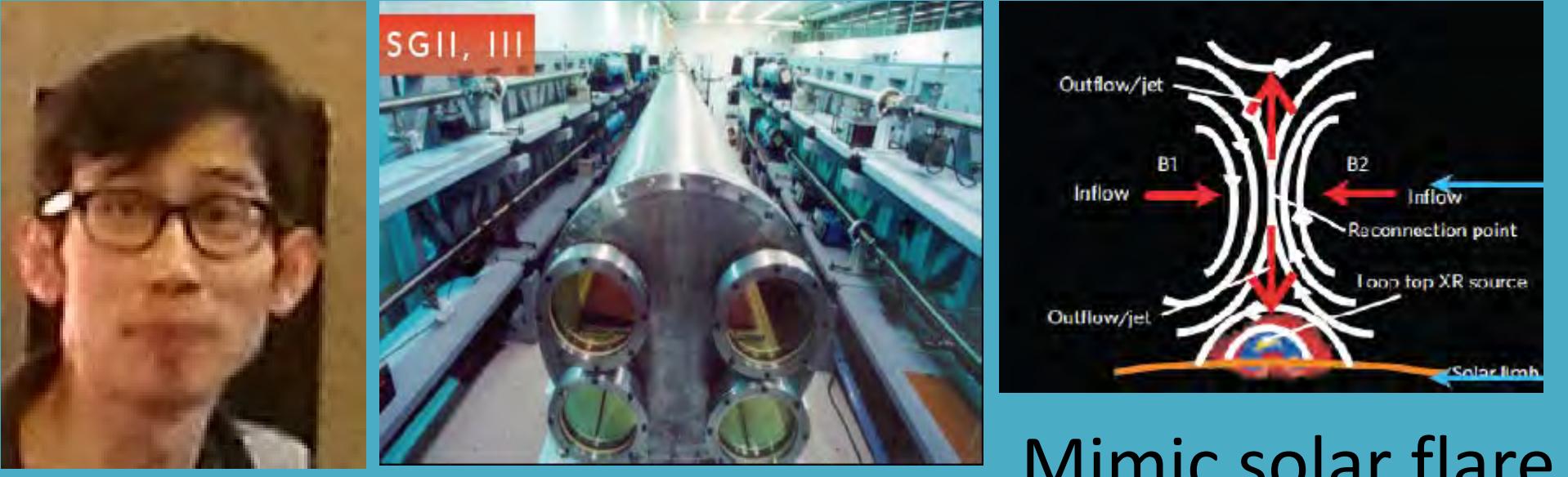


# 1 PW laser → reconnection



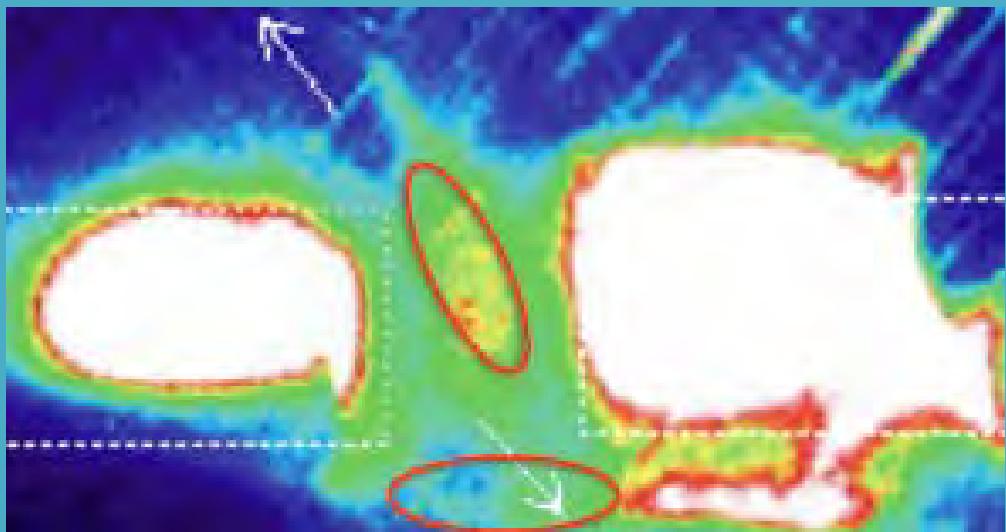
B. R. Lee





J. Y. Zhong

Mimic solar flare

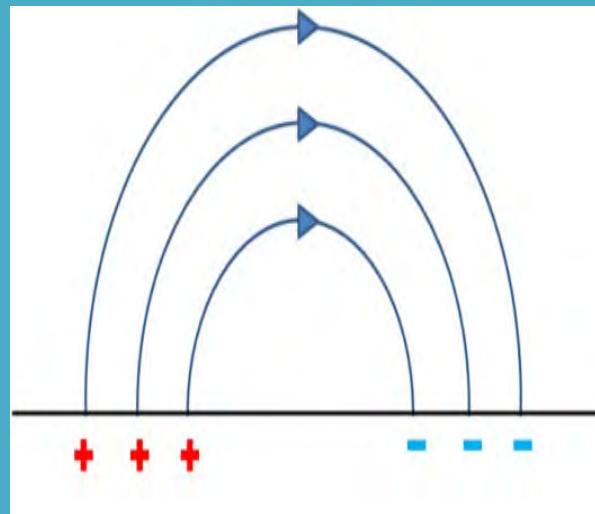


Low  $\beta$  & high  $\beta$

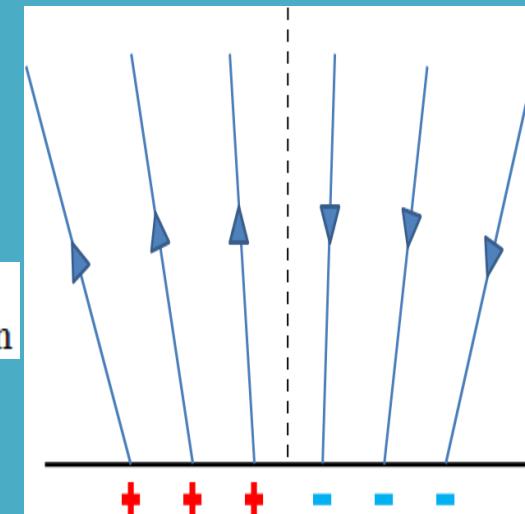
# Challenging the orthodox



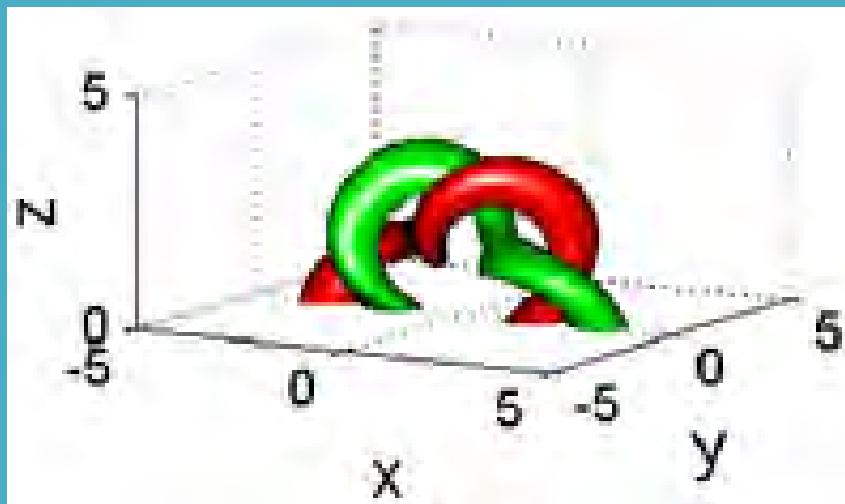
# Aly-Sturrock Theorem



$$E_{\text{fff}} \leq E_{\text{open}}$$



GS Choe



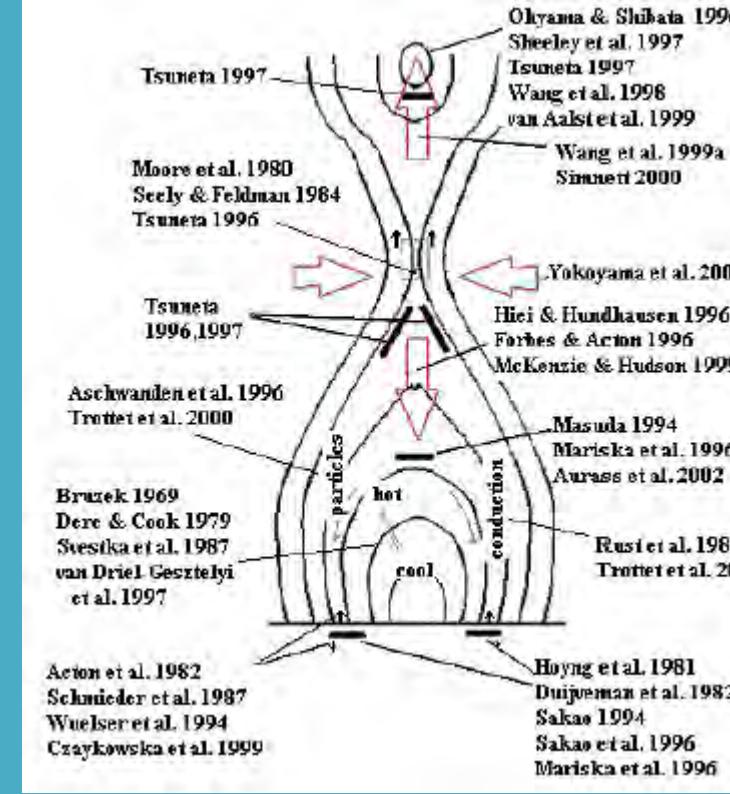
$$E_{\text{force-free}} > E_{\text{open}}$$



D. Melrose



Y. Bi



Standard reconnection model

Electromotive force (EMF)  
is missing, which is  
important.

Welcome back  
to China!



P. F. Chen

NANJING UNIVERSITY