

High-frequency electrostatic waves driven by electron crescents in the electron diffusion region

Wenya Li¹, Daniel B. Graham², Yu. V. Khotyaintsev², B. B. Tang¹, A. Vaivads³, M. Andre²,
K. Min⁴, K. Liu⁵, K. Fujimoto⁶, P.-A. Lindqvist³, K. Dokgo⁷, J. L. Burch⁷, C. Wang¹.

¹State Key Laboratory of Space Weather, National Space Science Center, Chinese Academy of Sciences, Beijing, China, ²Swedish Institute of Space Physics, Uppsala, Sweden, ³Division of Space and Plasma Physics, School of Electrical Engineering and Computer Science, KTH Royal Institute of Technology, Stockholm, Sweden, ⁴Department of Astronomy and Space Science, Chungnam National University, Daejeon, Republic of Korea, ⁵Department of Earth and Space Sciences, Southern University of Science and Technology, Shenzhen, China, ⁶School of Space and Environment, Beihang University, Beijing, China, ⁷Southwest Research Institute, San Antonio, Texas, USA.

e-mail (speaker): wyli@spaceweather.ac.cn

Magnetic reconnection is a fundamental and universal process, which transfers energy stored in the magnetic field to kinetic energy of charged particles¹. Magnetic reconnection powers eruptive processes in space and laboratory plasmas. Earth's magnetosphere provides a unique environment to study magnetic reconnection by analyzing in-situ spacecraft measurements. NASA's Magnetospheric Multiscale (MMS) mission² was designed to resolve the particles and fields at electron scales. The goal of MMS is to investigate the electron diffusion region (EDR), which is the core region of reconnection where the magnetic field lines break and reconnect. Reconnection at the dayside magnetopause is asymmetric due to the large plasma and magnetic field differences between the magnetosheath and the magnetosphere, while the reconnection in the magnetotail is mostly symmetric. Waves are suggested to generate anomalous resistivity and plasma diffusion, potentially enabling magnetic fields to break and reconnect. Various types of waves produced by reconnection have been reported outside of EDRs³.

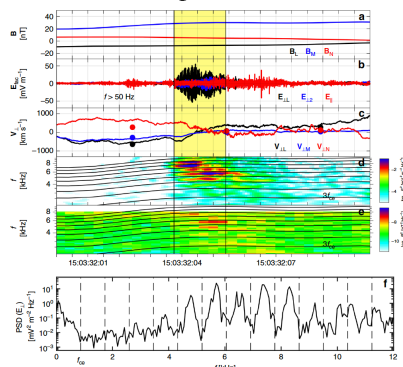


Fig. 1. Electron Bernstein waves observed by MMS⁴

The Magnetospheric Multiscale (MMS) spacecraft have encountered tens of electron diffusion regions (EDRs) of asymmetric magnetic reconnection at the Earth's

magnetopause and symmetric reconnection in the magnetotail. Electron crescent-shaped distributions are extensively observed in and near the electron diffusion and drive two types of electrostatic waves. One is the electron Bernstein waves (EBWs)⁴ and the other is upper-hybrid waves (UHWs)⁵⁻⁷. The EBWs are observed at the electron-scale boundary of the Hall current reversal in the outflow region of the magnetopause reconnection, while the UHWs are observed on the inflow side of EDRs of both dayside magnetopause and nightside magnetotail reconnections. The amplitudes of those high-frequency electrostatic waves are sufficiently large to thermalize and diffuse electrons. The strong interaction between the EBWs and UHWs can change the electron pressure tensor around the EDR and modify the balance of the magnetic reconnection electric field.

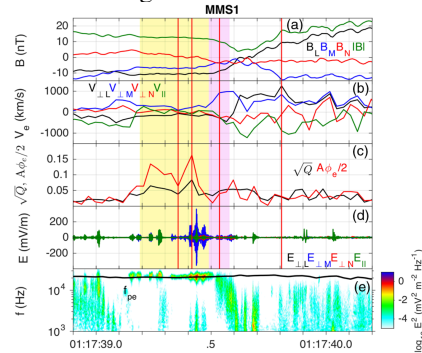


Fig. 2. Upper-hybrid waves observed by MMS⁵

References

- [1] Yamada *et al.*, Rev. Mod. Phys. 82, 603-664 (2010).
- [2] Burch *et al.*, Space Sci. Rev., 199, 5-21 (2016).
- [3] Khotyaintsev *et al.*, Front. Astron. Space Sci., 6:70 (2020).
- [4] Li *et al.*, Nature Communications, 11, 141 (2020).
- [5] Graham *et al.*, Phys. Rev. Lett., 119, 025101 (2017).
- [6] Burch *et al.*, Geophys. Res. Lett., 46 (2019).
- [7] Dokgo *et al.*, Geophys. Res. Lett., 47 (2020).