



Force-free magnetic island coalescence instability and Shear flow effects

Jagannath Mahapatra^{1,2}, Rajaraman Ganesh^{1,2}, and Abhijit Sen^{1,2}

¹Institute for Plasma Research, ²Homi Bhabha National Institute

e-mail (speaker): jagannath.mahapatra@ipr.res.in

Magnetic reconnection (MR) is responsible for some of the most energetic/eruptive events observed in space and laboratory plasmas, for example solar flares, coronal mass ejections, magnetospheric aurora formation, sawtooth crash in tokamaks, etc. during which the global magnetic energy is converted to kinetic energy, thermal energy and non-thermal particle acceleration [1]. The micro-scale current sheets (CSs) are the main site of MR and hence, it is a multi-scale phenomenon. Different plasma parameters around the CS controls the MR in many ways, for example, the plasma collisionality affects the CS dimension, the anti-parallel magnetic field profiles control the reconnection rate [2], shear flows controls the plasma upstream and downstream flow and CS orientation [2,3,5] etc.

Recently, using a visco-resistive incompressible MHD model [2], the effect of in-plane shear flow on 2D magnetic island coalescence and associated reconnection parameters have been investigated, both qualitatively [2] and quantitatively [3]. A new scaling law of reconnection rate with shear flow amplitude, different from that of Harris current sheet setup [5] has been reported. Moreover, the shear flow generated MHD-Kelvin-Helmholtz-instability modes are found to be suppressed by the coalescence instability. In the present work, we investigate the coalescence instability using a 2.5D force-free magnetic island equilibrium. A comparison of reconnection dynamics 2D, regular, non-force-free and a 2.5D force-free magnetic island equilibrium will be discussed. Further, the effect of shear flows on both type of initial profiles, a comparison with the incompressible results in the presence of a shear flow will be presented.

References :

- [1] E. Priest and T. Forbes, *The Astronomy and Astrophysics Review* 10, 313 (2002).
- [2] J. Mahapatra, A. Bokshi, R. Ganesh, and A. Sen, *Phys. Plasmas* 28, 072103 (2021).
- [3] J. Mahapatra, R. Ganesh, and A. Sen, *Phys. Plasmas* 29, 112107 (2022).
- [4] G. Murtas, A. Hillier, and B. Snow, *Phys. Plasmas* 28, 032901 (2021).
- [5] P. Cassak, *Phys. Plasmas* 18, 072106 (2011).