

Electron physics and global processes in the heliosphere

M.E. Innocenti¹, A. Micera¹, D. Verscharen², E. Boella³, A. Tenerani⁴, M. Velli⁵

¹Department of Physics, Ruhr University Bochum, Germany ²Mullard Space Science Laboratory, University College London, UK ³Department of Physics, University of Lancaster, UK ⁴Department of Physics, University of Texas at Austin, US ⁵Department of Earth, Planetary, and Space Sciences, University of California, Los Angeles, US
e-mail (speaker): mariaelena.innocenti@rub.de

Electron scales in the heliosphere are orders of magnitude smaller than global heliospheric scales, both in space and in time [1]. This notwithstanding, electron physics is fundamental in shaping global heliospheric processes.

The fundamental non-locality of electron physics in the heliosphere is highlighted by a number of recent works and observations.

The recently observed sunward deficit in the electron distribution is considered a signature of the interplanetary potential that traps electrons over global scales. The interplanetary potential traps the electrons that leave the Sun with kinetic energy below a cut-off value: these are the electrons that return towards the Sun. The electrons giving rise to the deficit are the ones with energy above the cut-off, which do not return [2,3]. The sunward deficit can create the conditions for an electron-scale microinstability, the whistler heat flux instability [4], which contributes to further modify the electron velocity distribution.

Another example of the fundamental role of electron physics in global heliospheric dynamics is given by heat flux regulation by electron scale microinstabilities. That kinetic instabilities should contribute to heat-flux regulation has been demonstrated decades ago with Ulysses observations [5]. Recent Parker Solar Probe and Solar Orbiter observations, e.g. [6], together with first-principle Particle-In-Cell simulations, e.g. [7], have shown a directly link between electron microinstabilities and scatter of strahl electrons, which results in reduction of the heat flux. Interestingly, also here global and electron-scale processes are related non-trivially: it has been demonstrated via semi-implicit expanding box Particle-In-Cell simulations [8] that plasma expansion can trigger electron-scale instabilities, which in turn contribute to heat flux regulation [9, 10].

In this talk, we will illustrate the above-mentioned links between electron scale and global heliospheric dynamics with the support of Particle In Cell simulations, both “standard” and Expanding Box.

References

- [1] Verscharen, D., Klein, K. G., & Maruca, B. A. (2019). The multi-scale nature of the solar wind. *Living Reviews in Solar Physics*, 16(1), 5.
[2] Berčič, L., Maksimović, M., Halekas, J. S., Landi, S., Owen, C. J., Verscharen, D., ... & Stevens, M. L. (2021). Ambipolar electric field and potential in the solar wind

estimated from electron velocity distribution functions. *The Astrophysical Journal*, 921(1), 83.

[3] Halekas, J. S., Berčič, L., Whittlesey, P., Larson, D. E., Livi, R., Berthomier, M., ... & Pulupa, M. P. (2021). The Sunward Electron Deficit: A Telltale Sign of the Sun’s Electric Potential. *The Astrophysical Journal*, 916(1), 16.

[4] Verscharen, D., Chandran, B. D. G., Boella, E., Halekas, J., Innocenti, M. E., Jagarlamudi, V. K., ... & Whittlesey, P. L. (2023). Electron-driven instabilities in the solar wind. *Plasma Waves in Space Physics: Carrying On the Research Legacies of Peter Gary and Richard Thorne*, 16648714.

[5] Scime, E. E., Bame, S. J., Feldman, W. C., Gary, S. P., Phillips, J. L., & Balogh, A. (1994). Regulation of the solar wind electron heat flux from 1 to 5 AU: Ulysses observations. *Journal of Geophysical Research: Space Physics*, 99(A12), 23401-23410.

[6] Cattell, C., Breneman, A., Dombek, J., Short, B., Wygant, J., Halekas, J., ... & Pulupa, M. (2021). Parker solar probe evidence for scattering of electrons in the young solar wind by narrowband whistler-mode waves. *The Astrophysical journal letters*, 911(2), L29.

[7] Micera, A., Zhukov, A. N., López, R. A., Innocenti, M. E., Lazar, M., Boella, E., & Lapenta, G. (2020). Particle-in-cell simulation of whistler heat-flux instabilities in the solar wind: Heat-flux regulation and electron halo formation. *The Astrophysical Journal Letters*, 903(1), L23.

[8] Innocenti, M. E., Tenerani, A., & Velli, M. (2019). A Semi-implicit Particle-in-cell Expanding Box Model Code for Fully Kinetic Simulations of the Expanding Solar Wind Plasma. *The Astrophysical Journal*, 870(2), 66.

[9] Innocenti, M. E., Boella, E., Tenerani, A., & Velli, M. (2020). Collisionless heat flux regulation via the electron firehose instability in the presence of a core and suprathermal population in the expanding solar wind. *The Astrophysical Journal Letters*, 898(2), L41.

[10] Micera, A., Zhukov, A. N., López, R. A., Boella, E., Tenerani, A., Velli, M., ... & Innocenti, M. E. (2021). On the role of solar wind expansion as a source of whistler waves: Scattering of suprathermal electrons and heat flux regulation in the inner heliosphere. *The Astrophysical Journal*, 919(1), 42.